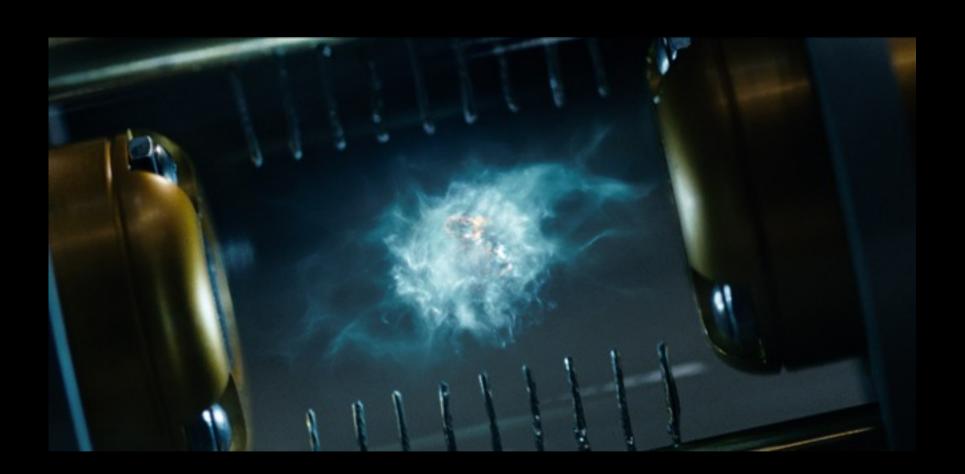
Neutrinos and the Case of the Missing Antimatter

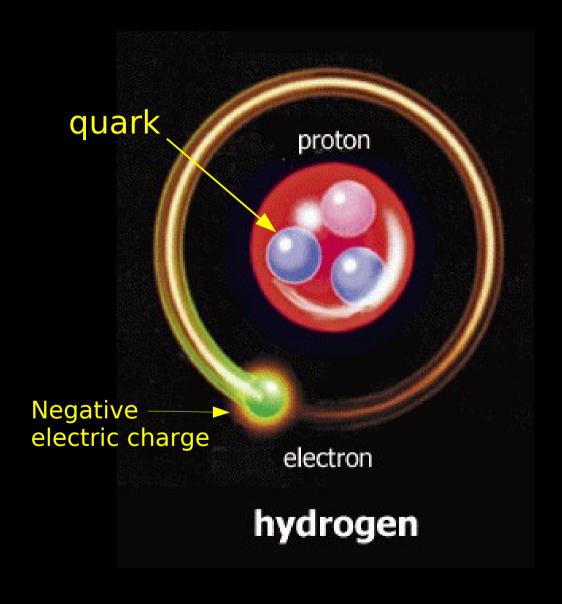


Not this...



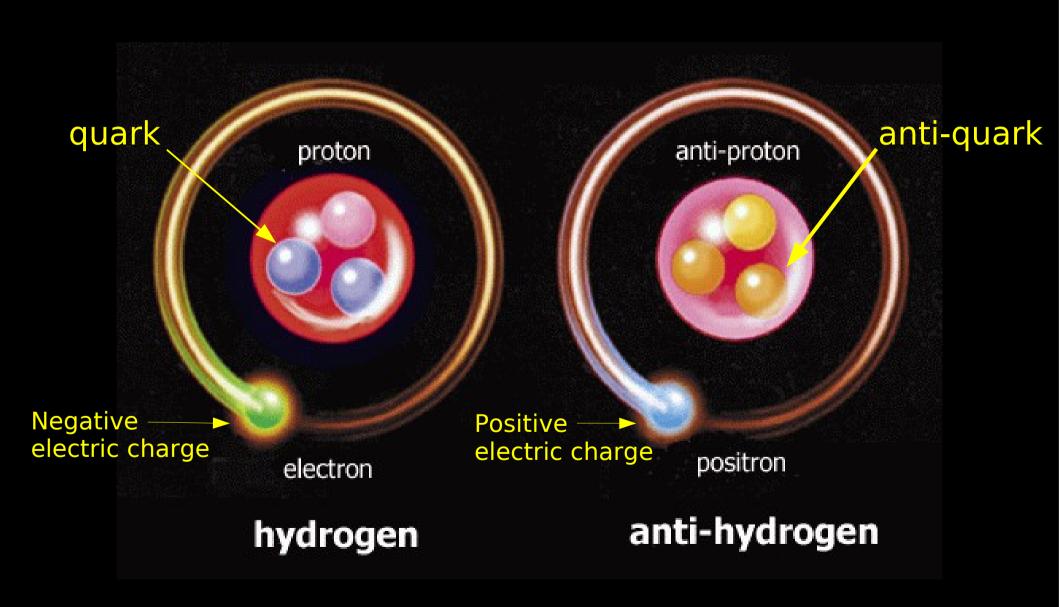
Angels and Demons, 2009

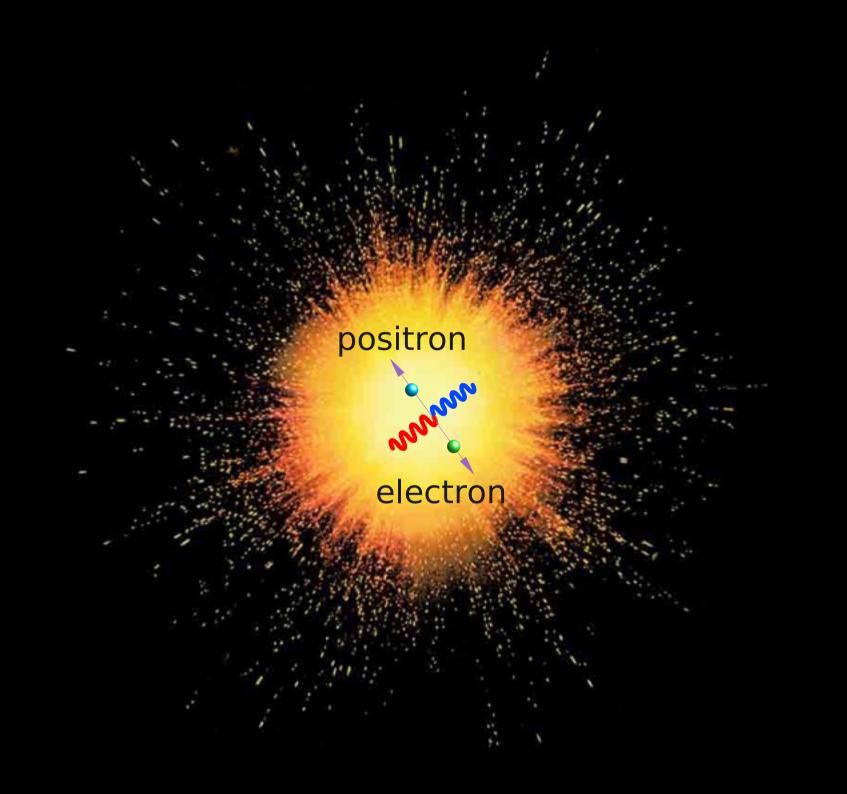
Matter



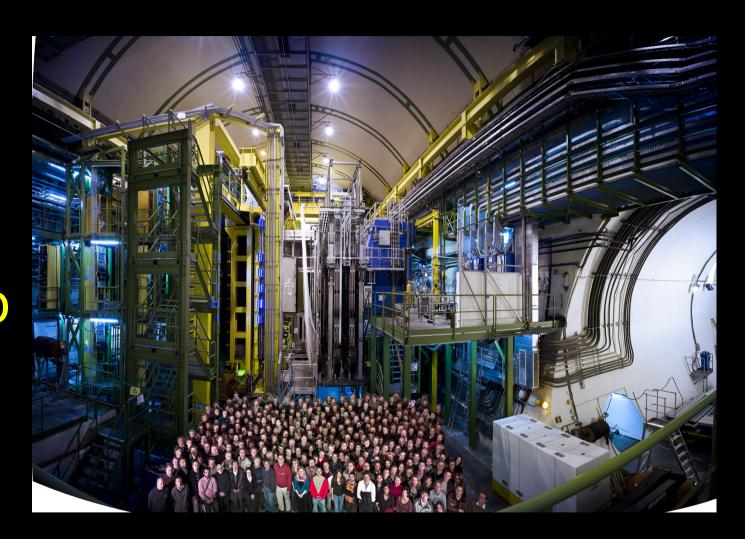
Matter

Anti-Matter





There is a difference between the physics of matter and antimatter. It's name is CP Violation



LHCb

The LHC will study this by looking differences between particles called B⁰ and B⁰ mesons

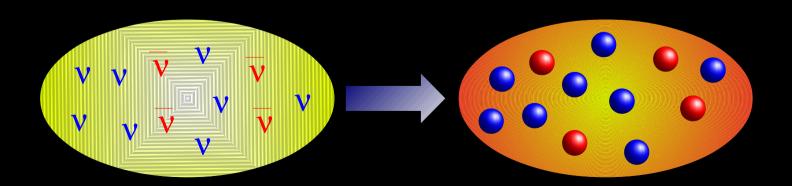
But.....

We might be looking in the wrong place....

The smallest, most insignificant (yet most common) particle in the cosmos may just hold the reason!

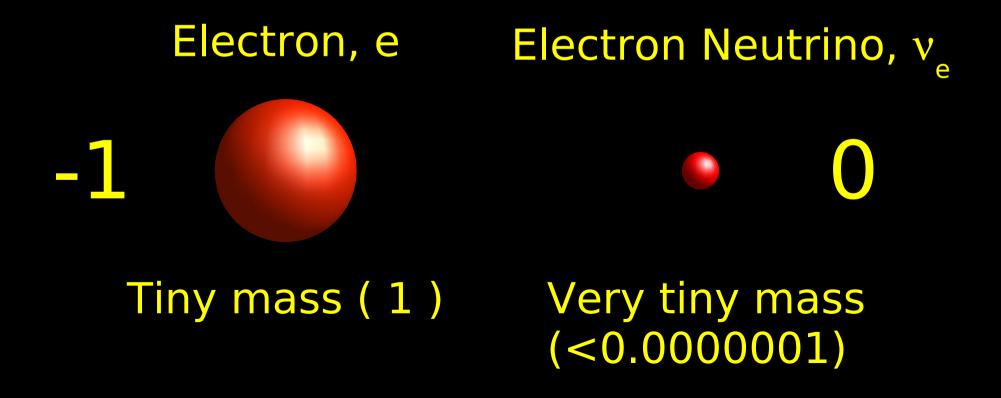
Matter-Antimatter Asymmetry

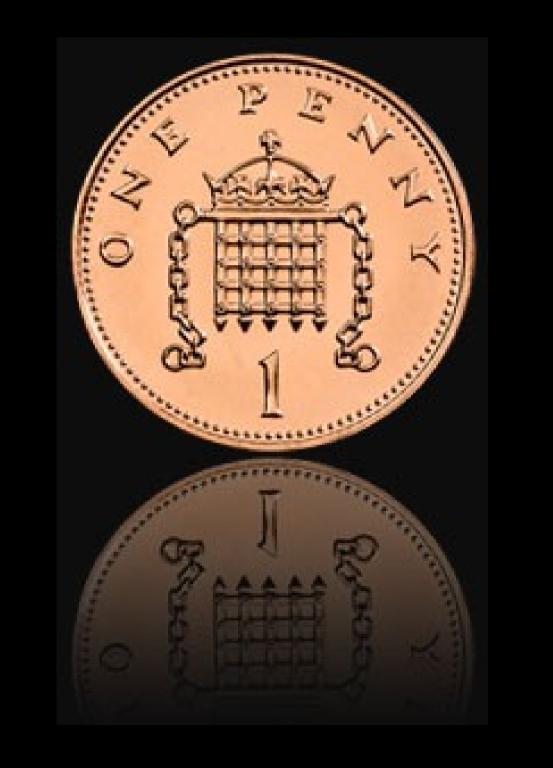
A idea called "Leptogenesis" suggests that the asymmetry we see between matter and antimatter could have been generated by an asymmetry between neutrinos and anti-neutrinos at the beginning of things.



So what is a neutrino?

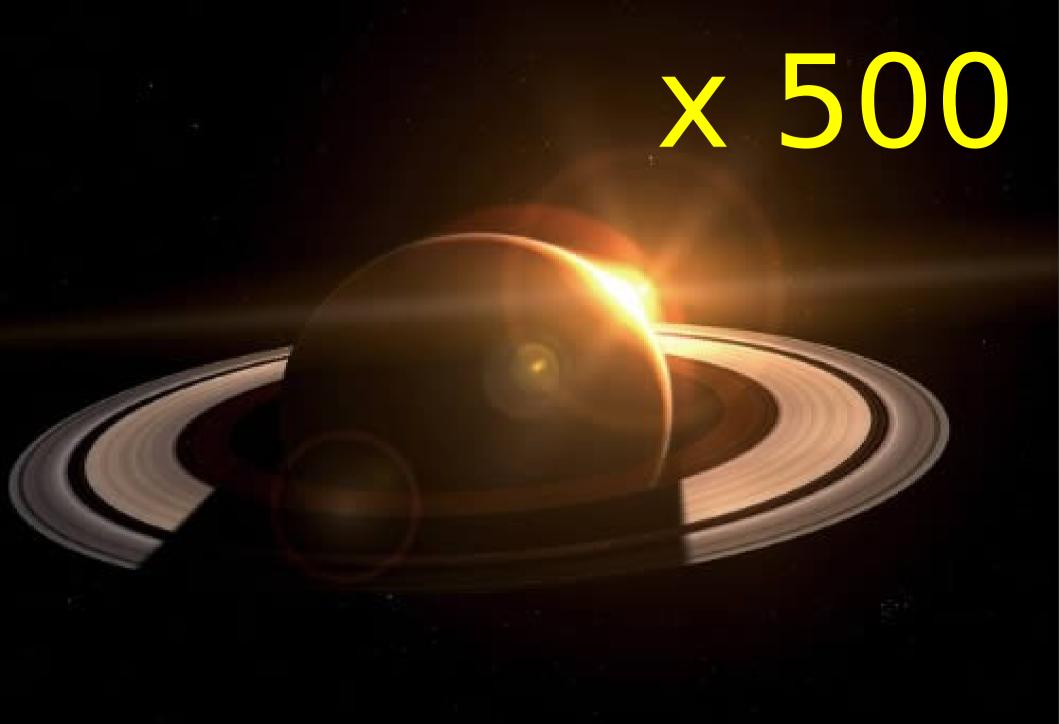
Neutrinos are the second most common particle in the universe. They are produced whenever something radioactively decays











Electron, e mass (1)



Electron Neutrino, ν_e

Muon, μ mass (200)



 $\begin{array}{c} & \text{Muon} \\ & \text{Neutrino, } \nu_{\mu} \end{array}$

Tau, τ mass (3500)



Tau Neutrino, ν_τ

3 <u>Lepton Flavours</u> + anti-leptons Electron Neutrino, v_e

Electron
Antineutrino, \overline{v}_{e}

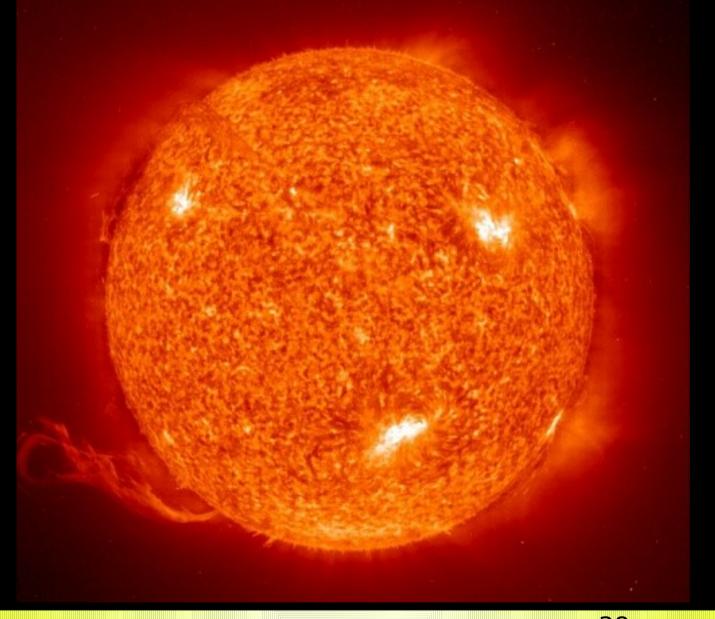
Muon Neutrino, ν_{μ}

Muon
Antineutrino, $\overline{\nu}_{\mu}$

Tau Neutrino, v_{τ}

Tau
Antineutrino, $\overline{\nu}_{\mu}$

3 neutrino Flavours

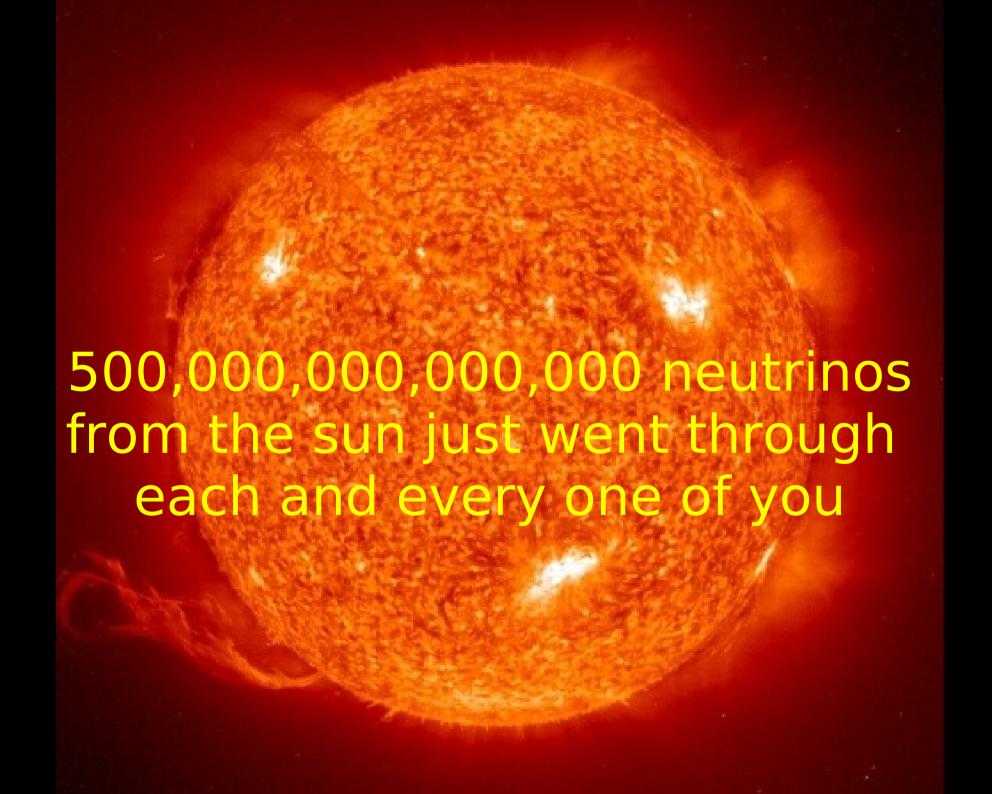


The sun generates about 2x10³⁸ neutrinos/s as byproducts of the fusion processes that make the star shine.

So why don't we notice?

v are almost ghosts. They interact extremely weakly with matter.

To a neutrino a planet is mostly empty space.



"The chances of a neutrino actually hitting something as it travels through all this howling emptiness are roughly comparable to that of dropping a ball bearing at random from a cruising 747 and hitting, say, an egg sandwich."

Douglas Adams-Mostly Harmless

Probability $\approx 5 \times 10^{-13}$





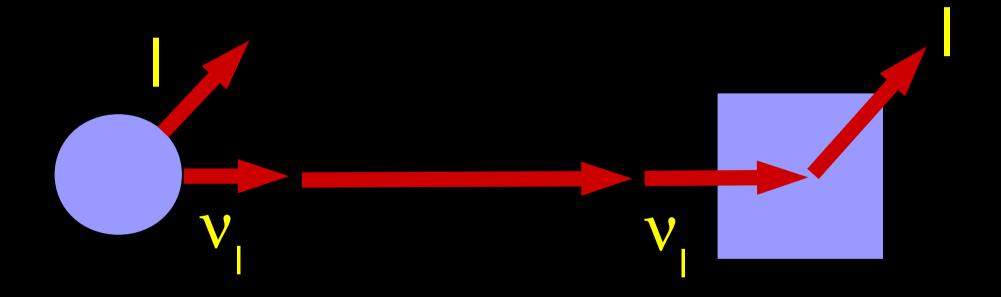




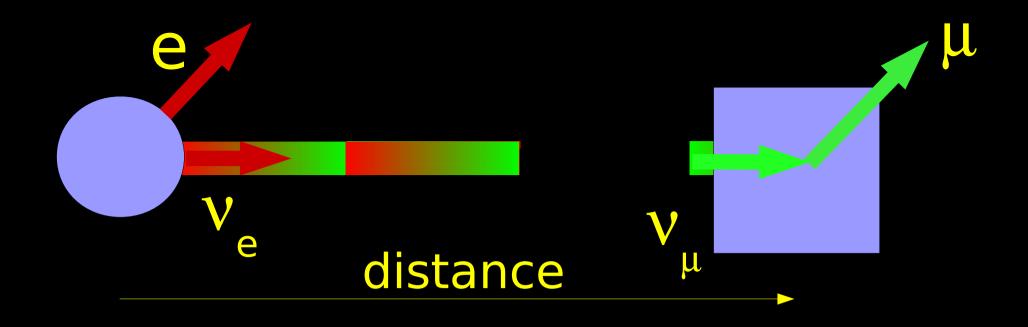
How do we use neutrinos to study CP violation?

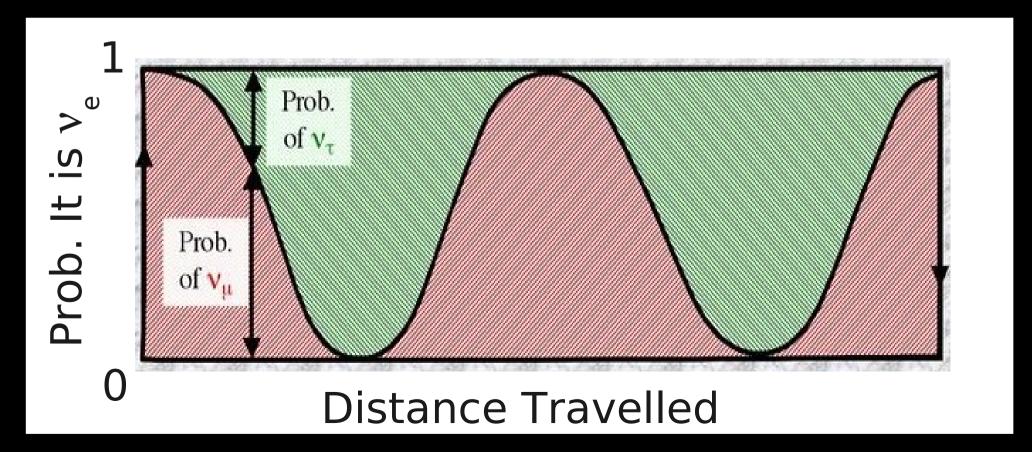
Neutrino Oscillations

THE discovery in neutrinos of the last 20 years



A typical neutrino experiment





T2K Experiment

Super Kamiokande



IPARC

295 km

Image € 2008 TerraMetrics
Image NASA

Image © 2008 Digital Earth Technology

***Google**

The Master Plan

Measure oscillations of neutrinos

Measure oscillations of anti-neutrinos

Difference between these two numbers indicates a difference between the physics of neutrino and anti-neutrinos

Can help show that leptogenesis works

Experiment is starting now....

But we still don't know much about the neutrino itself!

- •What is the mass of a neutrino?
- •Why are they so much lighter than all the other massive particles?
- •Are neutrinos the same as antineutrinos?
- Are neutrinos the reason we are here at all?

Economic Impacts

- 5% of jobs in UK are in physics-based sectors
 Gross added value from physics sector was estimated to be 70 billion pounds in 2005
 Synergy between PP projects and industry industry
- acquires added skills base for other applications
 Training 50% of PP PhDs go into other sectors
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 - Muon tomography in border security
 - Airport scanners
 - **Rock Imaging**
 - Cancer treatment using next gen cyclotrons proton therapy

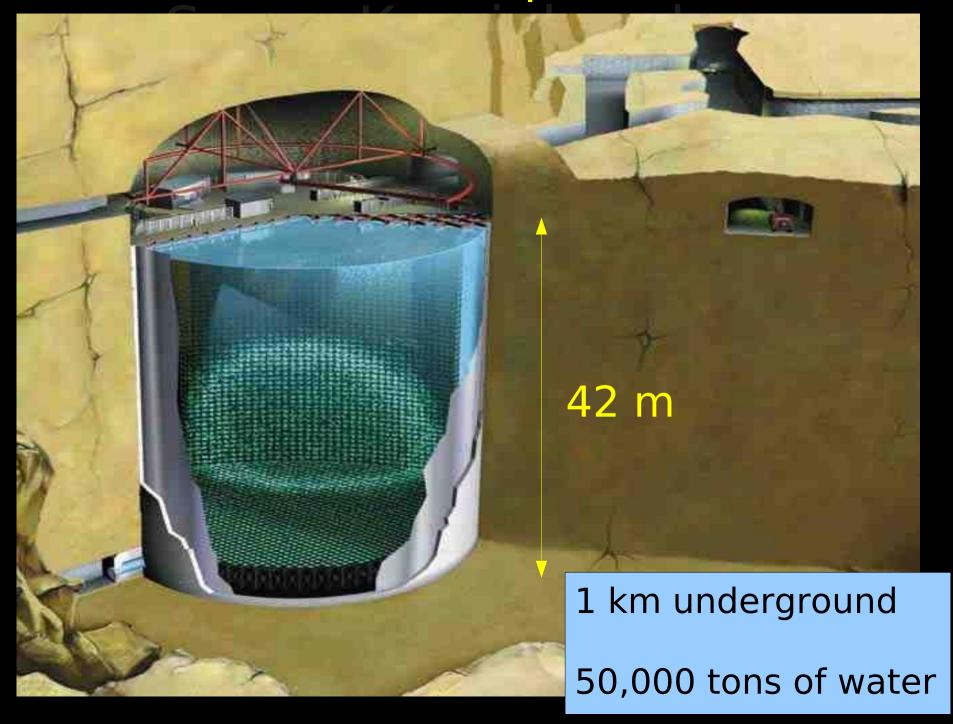
How do we exist?

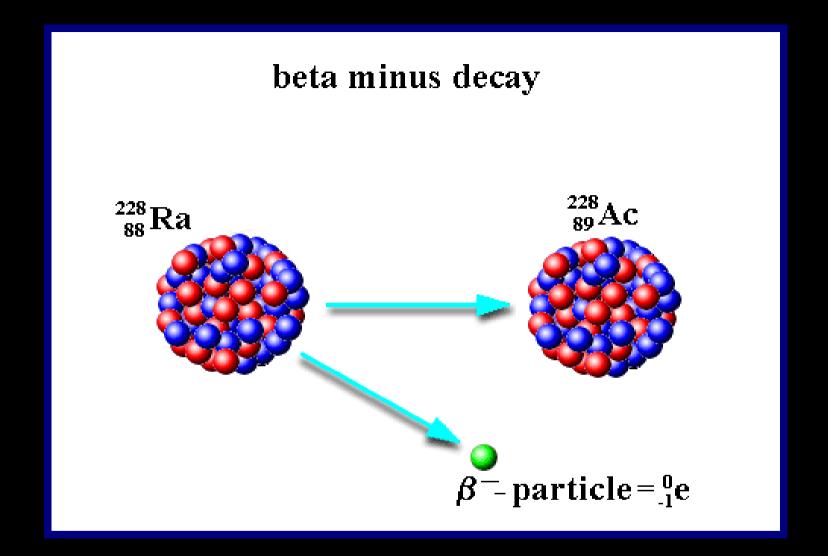
We don't know (yet) but we're working on it

The smallest, most insignificant (yet most common) particle in the cosmos may just hold the reason!



Super-Kamiokande





Energy(Ra) \neq Energy(Ac)+Energy(e)





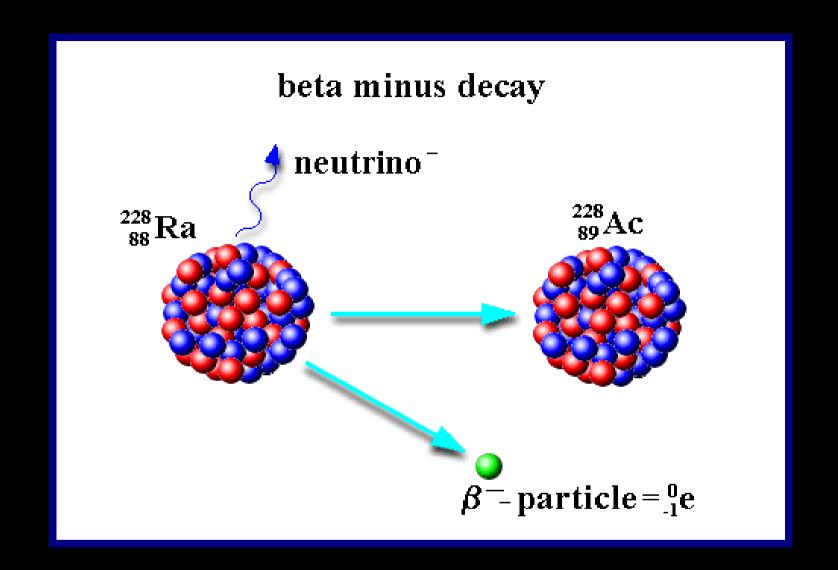
```
"Desperate remedy....."

"I do not dare publish this idea...."

"I admit my way out may look
improbable...."

"Weigh it and pass sentence...."
```

"You tell them. I'm off to a party"



Energy(Ra) = Energy(Ac)+Energy(e) + Energy(Neutrino)

What are neutrinos?

Electron, e

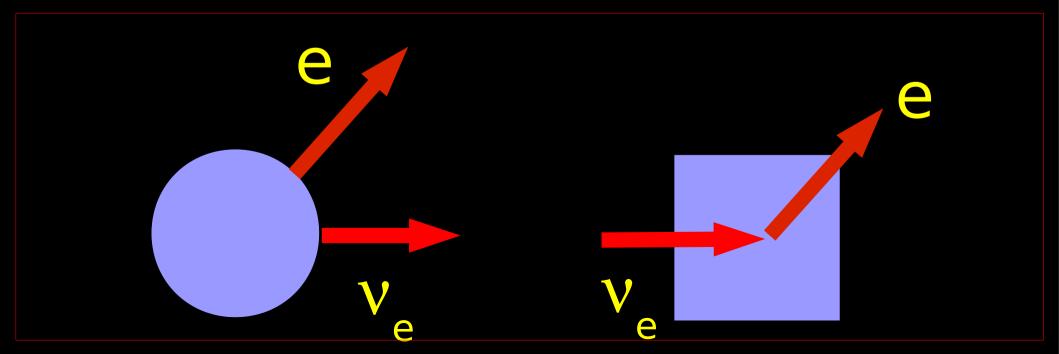
Electron Neutrino, v_e





Tiny mass (1) Very tiny mass

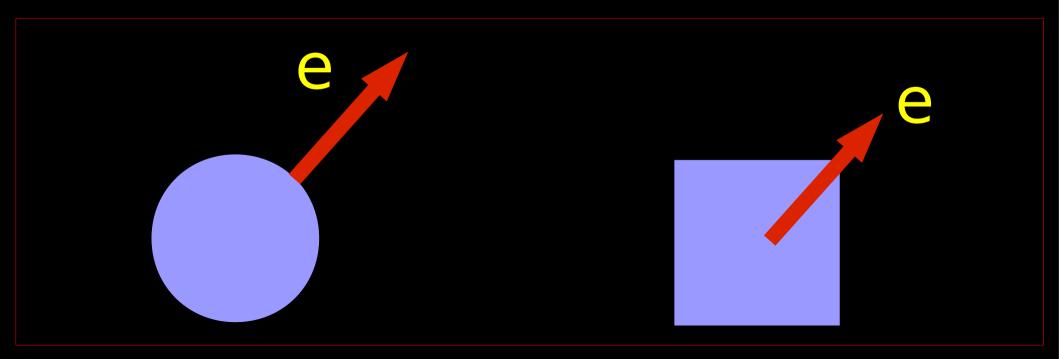
(<0.000001)

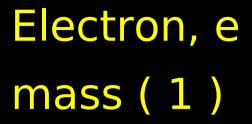


In experiments neutrinos are NEVER seen.

We can only detect them through the byproducts of their interactions with matter.

Type of the charged particle detected used to infer the type of incoming neutrino.







Electron Neutrino, v_e

Muon, μ mass (200)



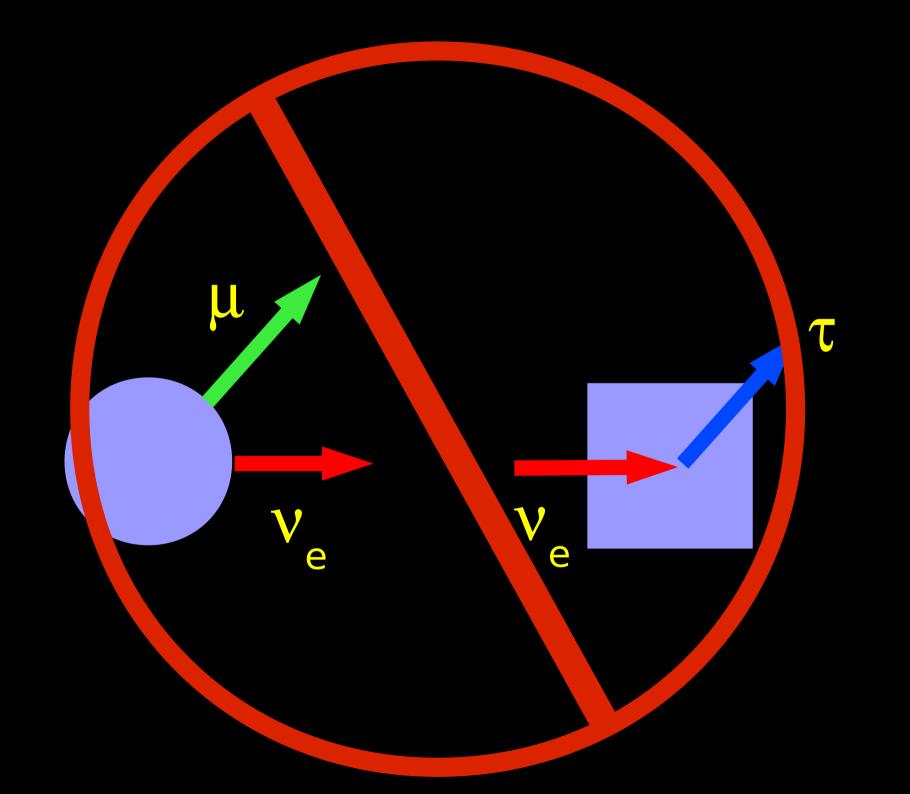
Muon Neutrino, ν_μ

Tau, τ mass (3500)



Tau Neutrino, ν_τ

3 Lepton Types



Positron, e⁺ mass (1)



Electron
Antineutrino, \overline{v}_{e}

Muon, μ^+ mass (200)



Muon Antineutrino, v

Tau, τ⁺
mass (3500)



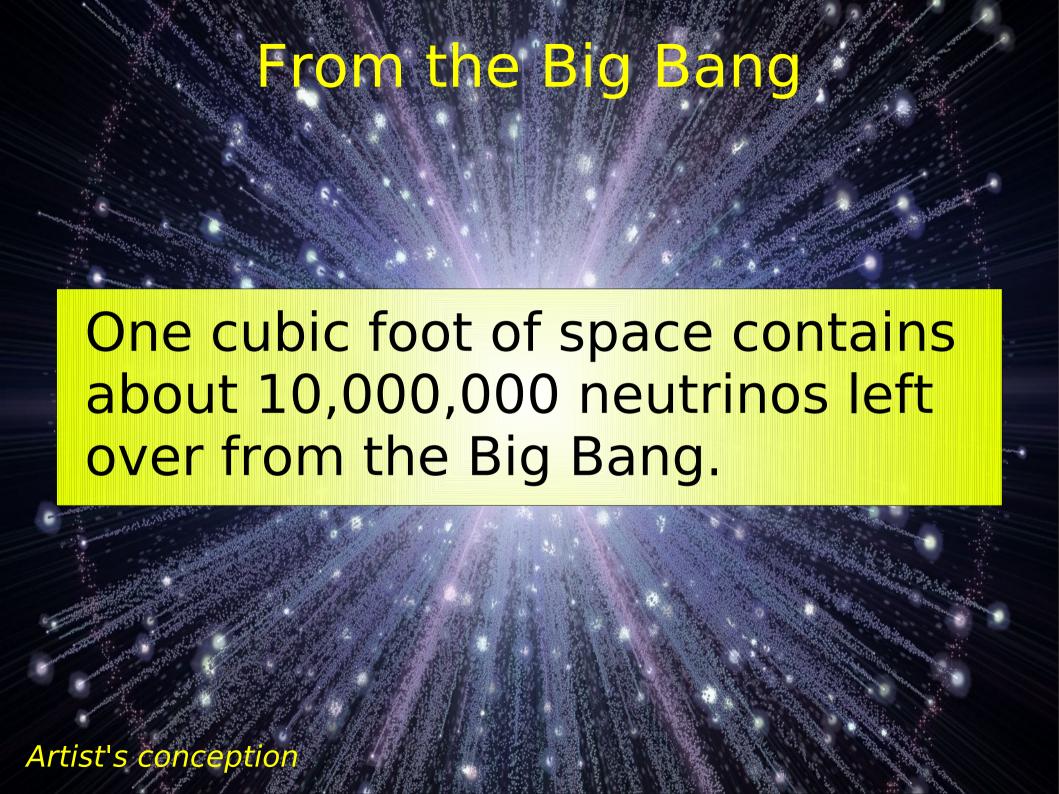
Antineutrino, $\overline{\nu}_{\mu}$

3 Antiparticles

Where do they come from?

Everywhere!



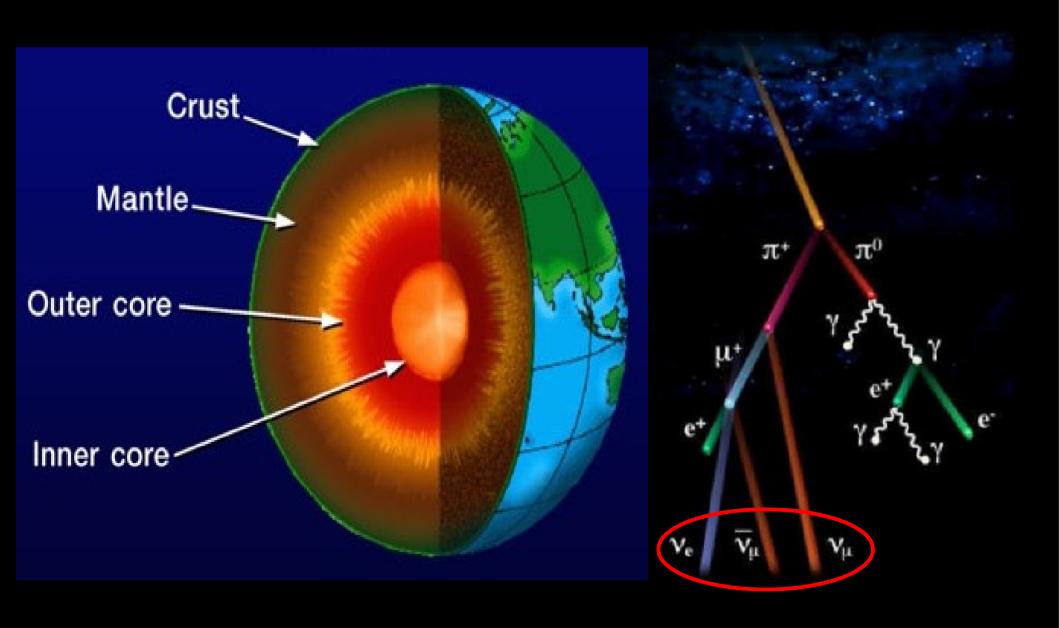




Supernovae created the heavy elements (us) and neutrinos may be responsible for the star exploding.



From The Earth





So why don't we notice?

v are almost ghosts. They interact extremely weakly with matter.

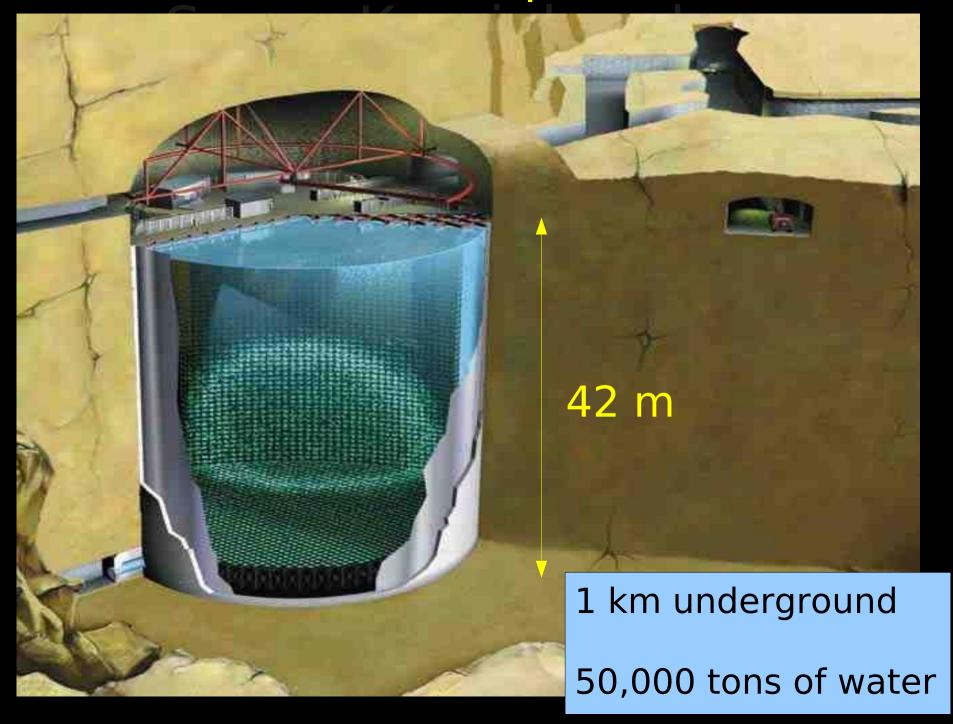
To a neutrino a planet is mostly empty space.

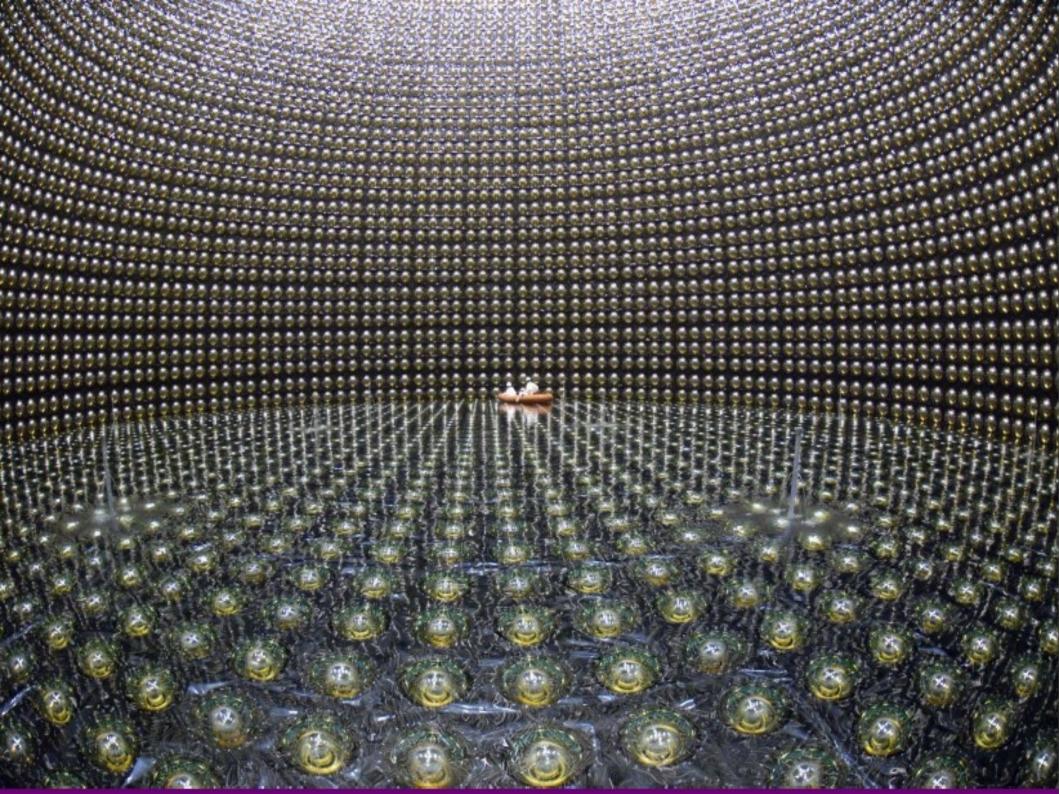
"The chances of a neutrino actually hitting something as it travels through all this howling emptiness are roughly comparable to that of dropping a ball bearing at random from a cruising 747 and hitting, say, an egg sandwich."

Douglas Adams



Super-Kamiokande



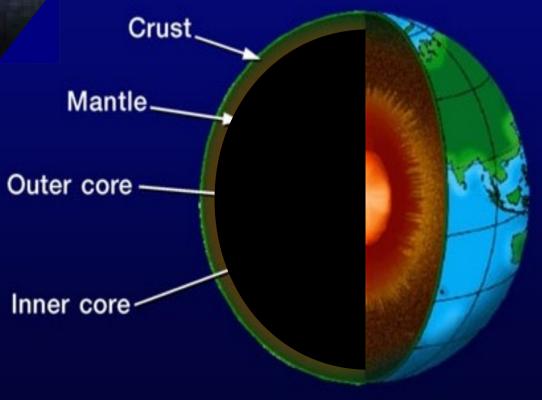


Why do we study them?

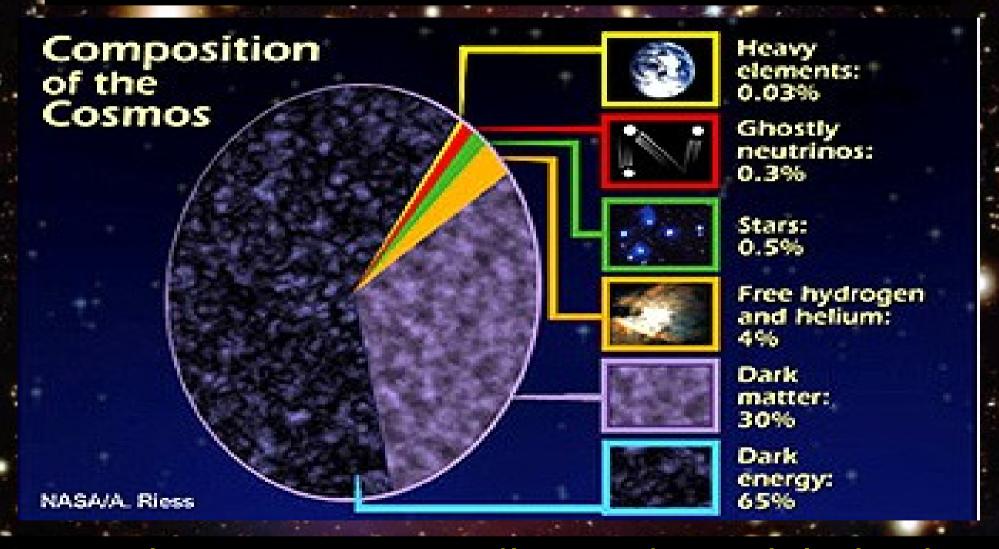


As Probes

Astrophysics
Geophysics
Cosmology
Particle Physics

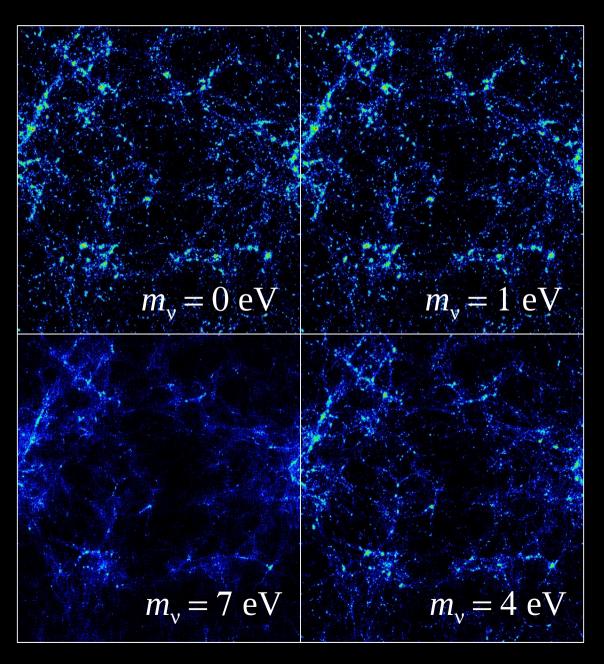


Amount of matter in Universe

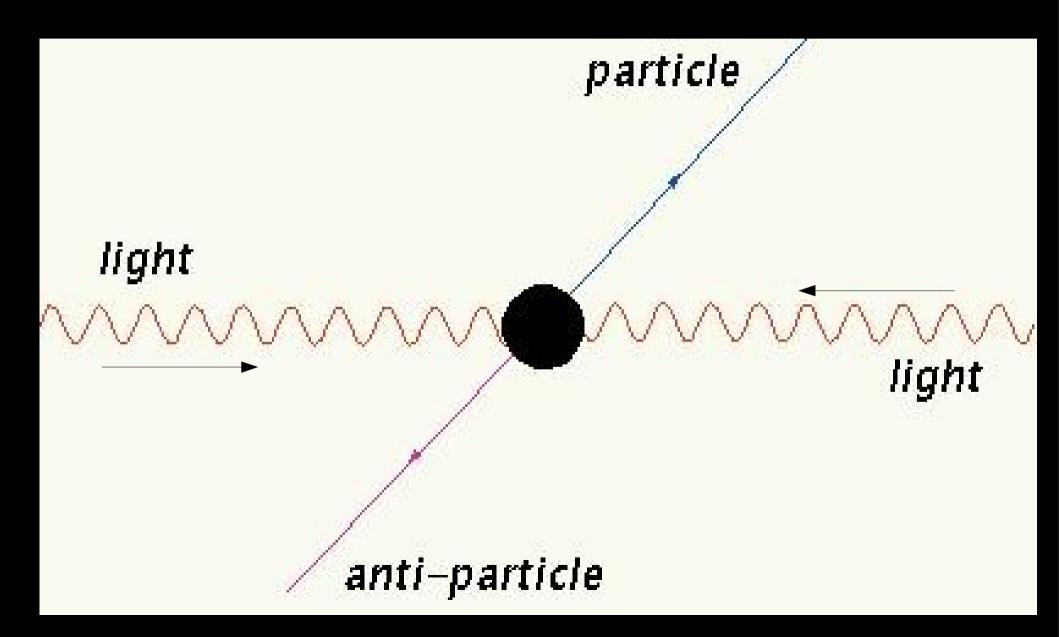


Second most common (known) particle in the universe

Universal Structure



Why is there more matter than antimatter?



Why is there more matter than antimatter?

10,000,000,001

10,000,000,000

Matter

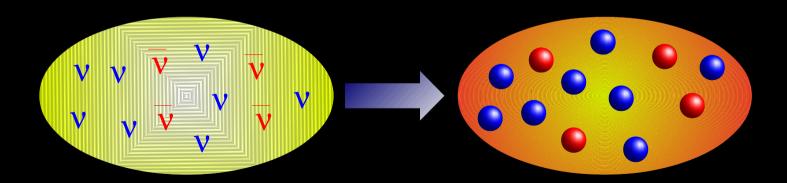
Antimatter

Matter-Antimatter Asymmetry

Q. Is there a difference between the physics of matter and antimatter?

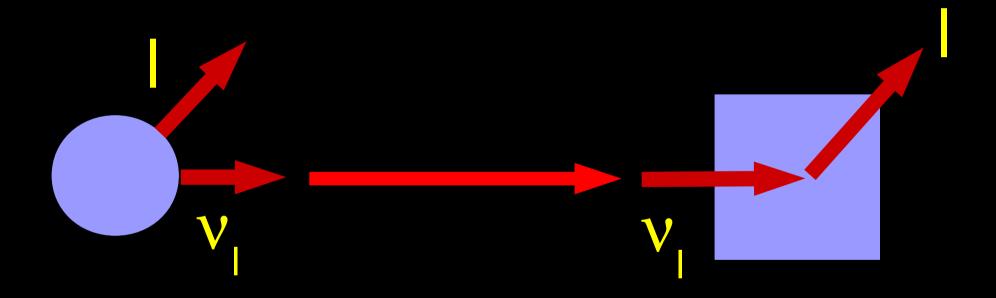
A.Yes there is.

We've never seen it in neutrinos, though.



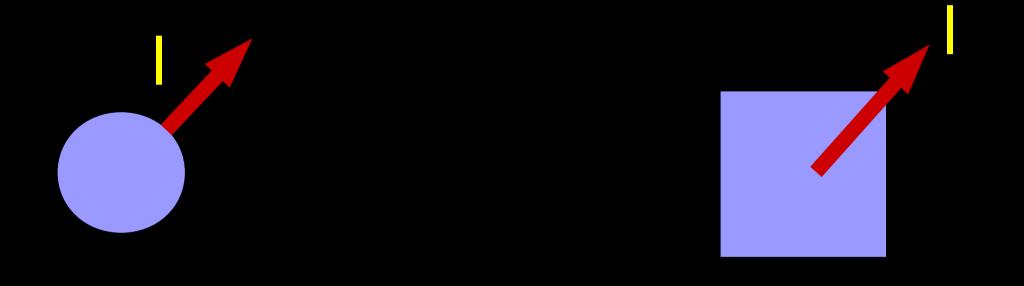
How to study this...?

THE discovery in neutrinos of the last 20 years



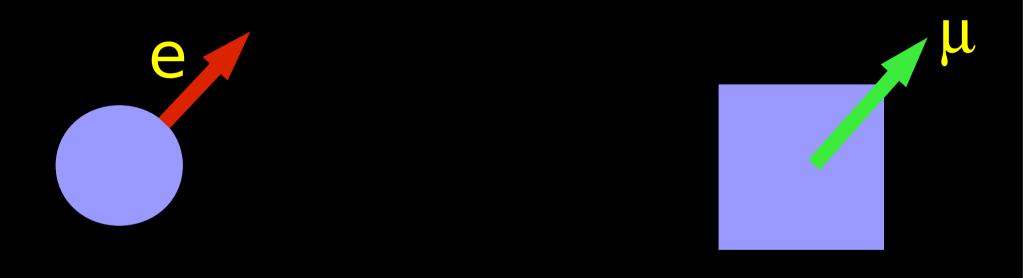
A typical neutrino experiment

THE discovery in neutrinos of the last 20 years



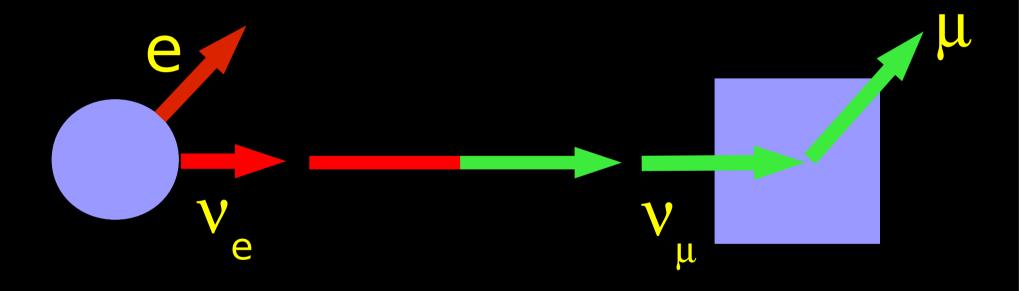
A typical neutrino experiment

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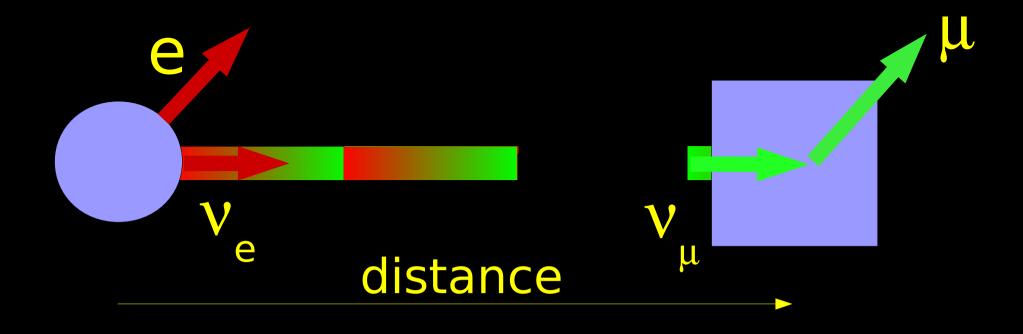


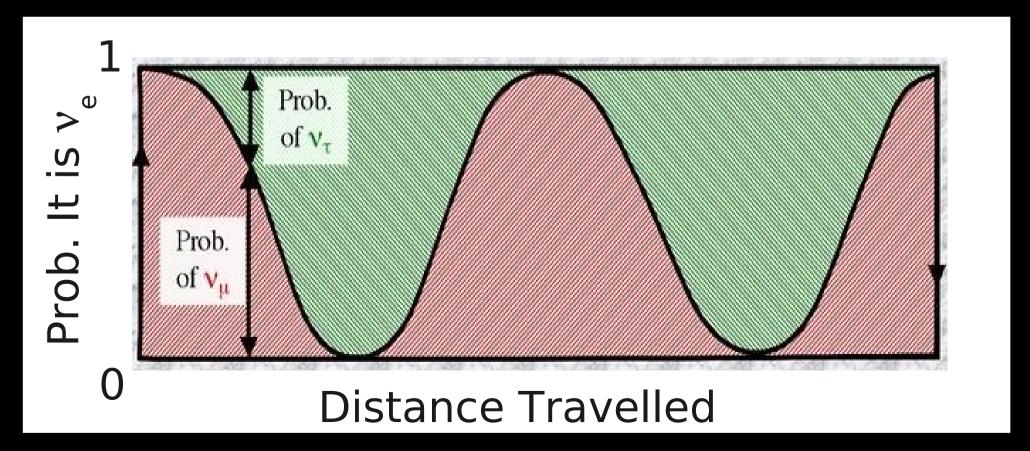
A typical neutrino experiment

THE discovery in neutrinos of the last 20 years



Neutrinos change flavour between source and detector!



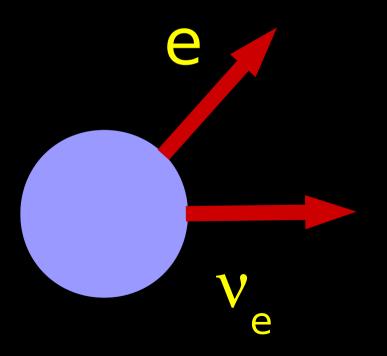


What the....?

Q. How can a v_e spontaneously turn into a v_{μ} ?

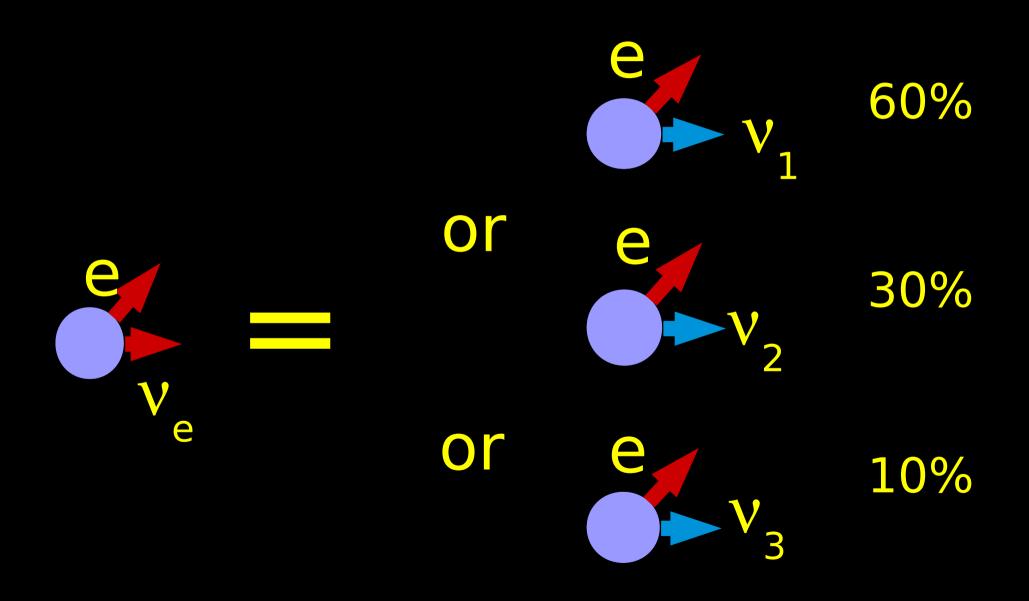
What the....?

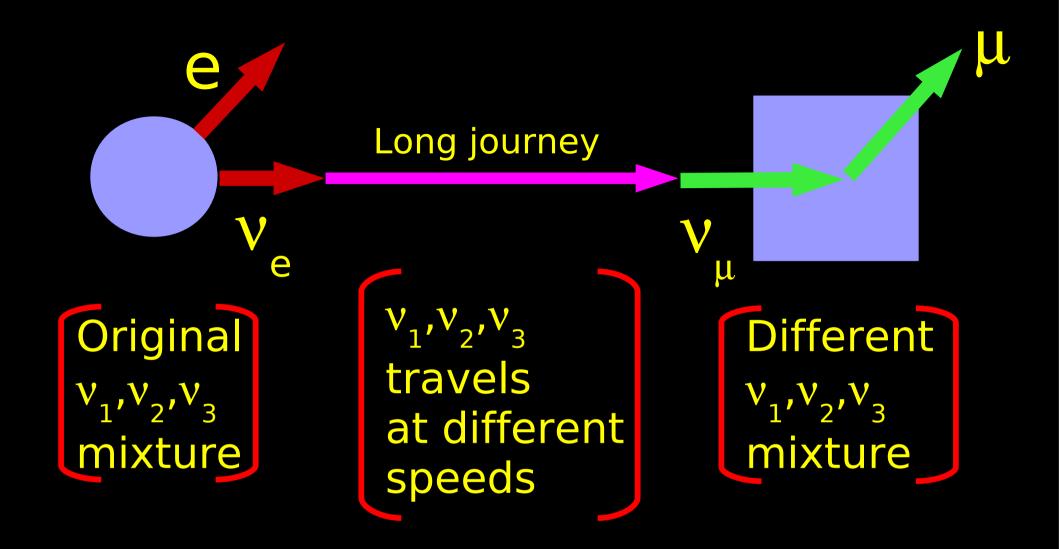
Q. How can a v_e spontaneously turn into a v_{μ} ? A. The v_e isn't *a* particle. It's three!



 $v_e \equiv$ "that thing which was always produced/detected with an electron"

Quantum Stuff





- •This can only happen if v_1, v_2, v_3 have different masses
- Only gives us differences in masses

T2K



SuperKamiokande

JPARC

295 km

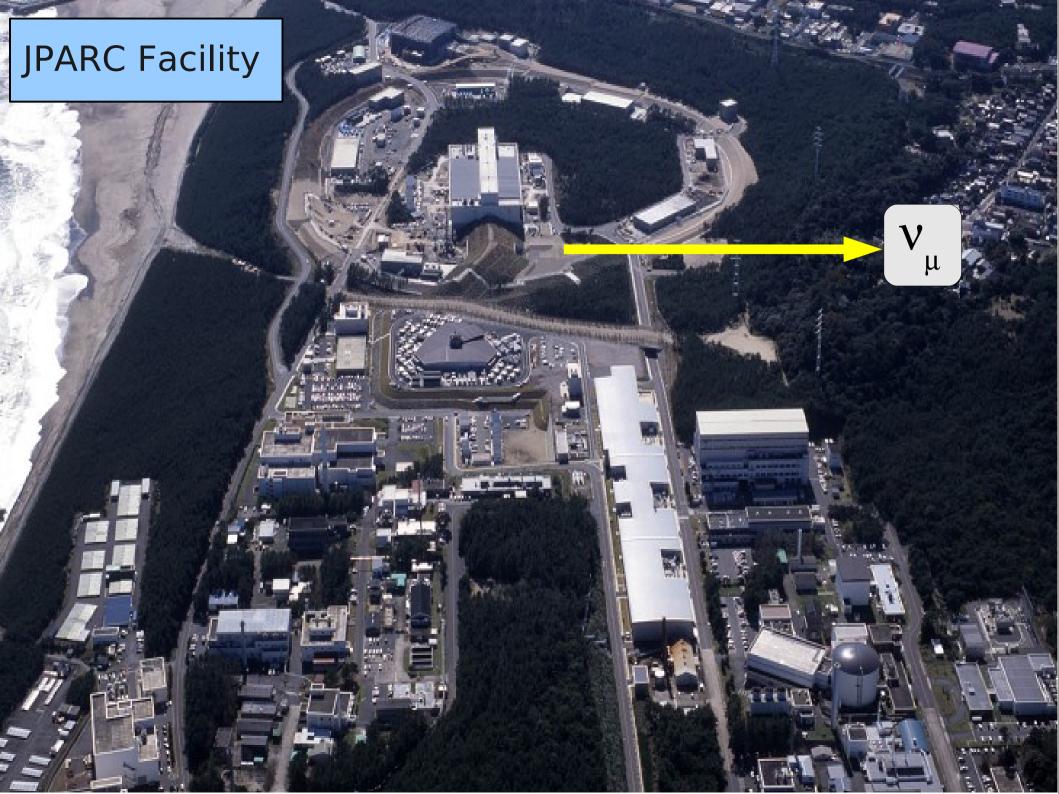
Image € 2008 TerraMetrics Image NASA

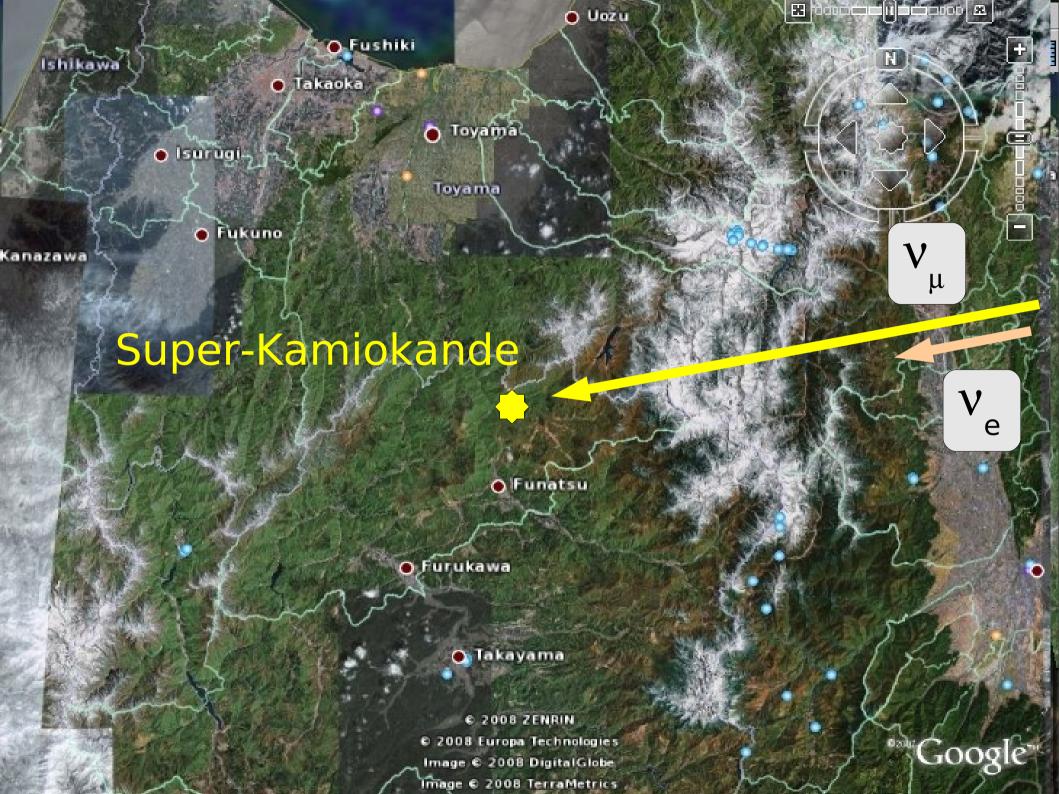
Image € 2008 Digital Earth Technology

*2007 Google™

Things we still don't know

- •How much do v_1, v_2 and v_3 weigh?
- •Why are they so much lighter than all the other massive particles?
- •Are neutrinos the same as antineutrinos?
- •Are neutrinos the reason we are here at all?





Economic Impacts

- •5% of jobs in UK are in physics-based sectors
- •Gross added value from physics sector was estimated to be 70 billion pounds in 2005
- •Synergy between PP projects and industry industry acquires added skills base for other applications
- •Training 50% of PP PhDs go into other sectors
 - Radioisotope production
 - Sensors for medical applications
 - High level computing for biological/climate modelling
 - Spin off tools for other science (e.g. DIAMOND)
 - Nuclear fusion research
 - Muon tomography in border security
 - Airport scanners
 - **Rock Imaging**
 - Cancer treatment using next gen cyclotrons
 - proton therapy

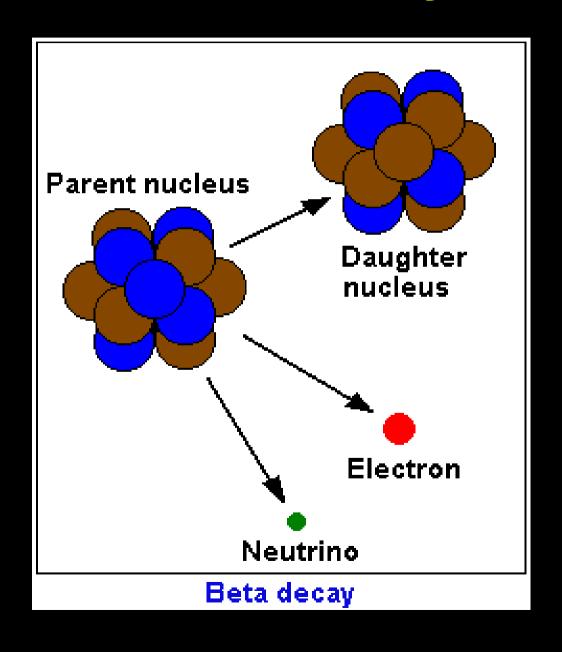
"...these kind of findings have implications that are not limited to the laboratory. They affect the whole of society — not only our economy, but our very view of life, our understanding of our relations with others, and our place in time."

Bill Clinton

From Radioactive Decay

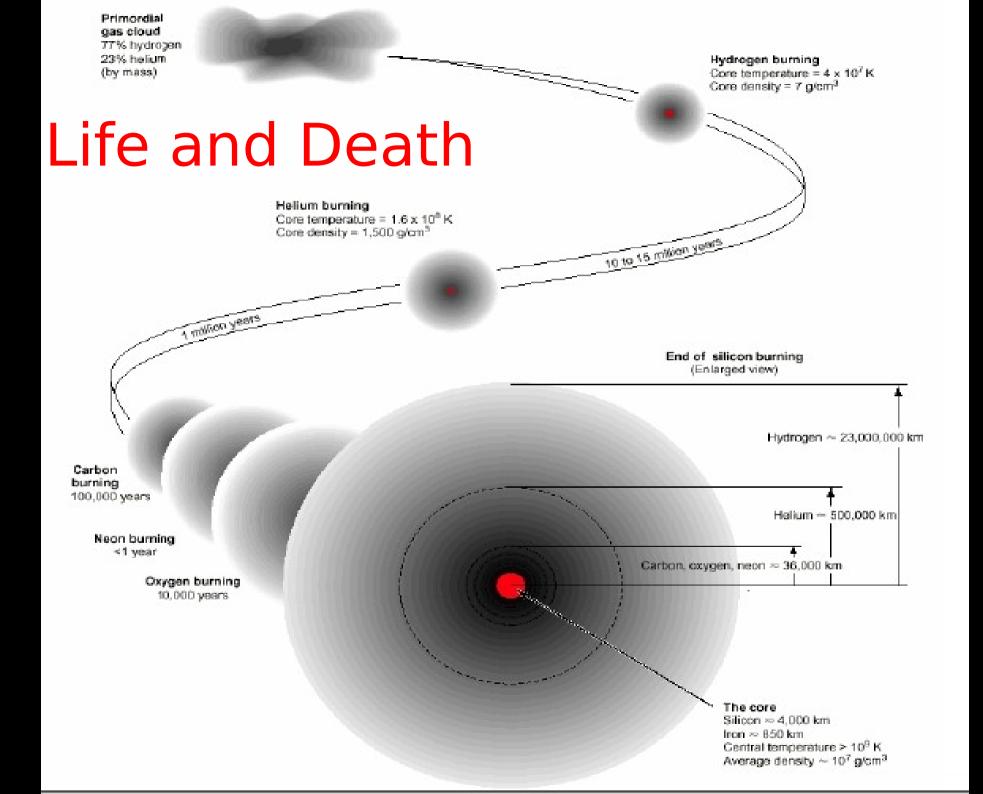
 $n \rightarrow p + e + v_e$

e.g. Decay of unstable nuclides in the core of the earth can tell us about its structure (Geoneutrinos)

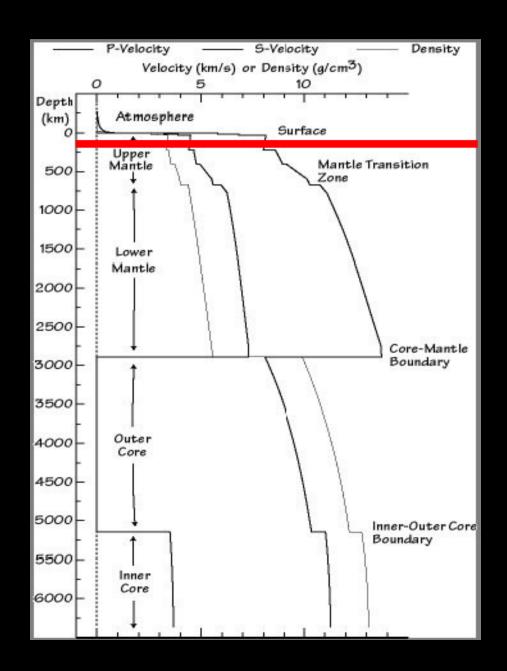




"Quarks. Neutrinos. Mesons. All those damn particles you can't see. <u>That's</u> what drove me to drink. But <u>now I</u> can <u>see</u> them!"



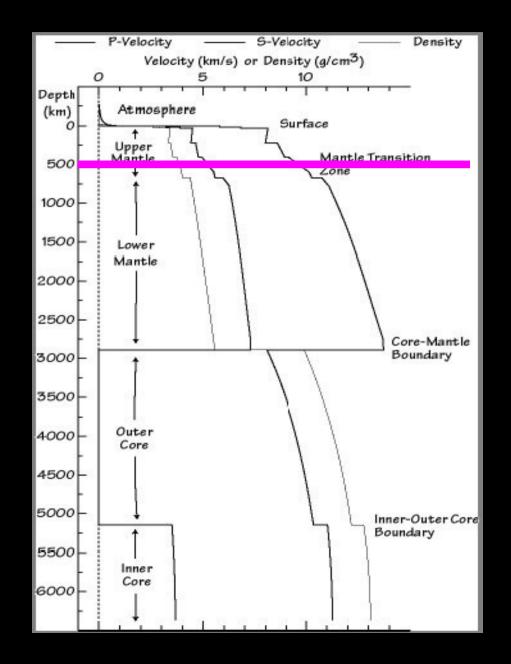
Geoneutrinos



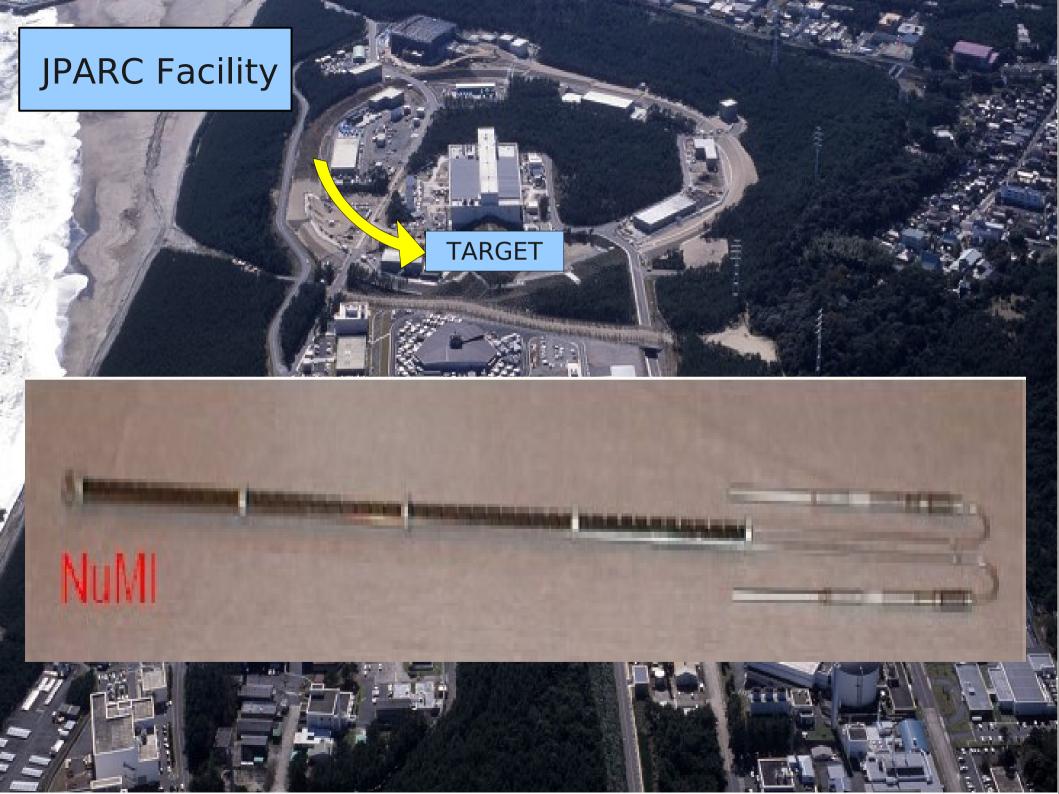
Models suggest A total heatflow of 19 TW from radioactive decay

A neutrino experiment in Japan measured 25 ± 20 TW

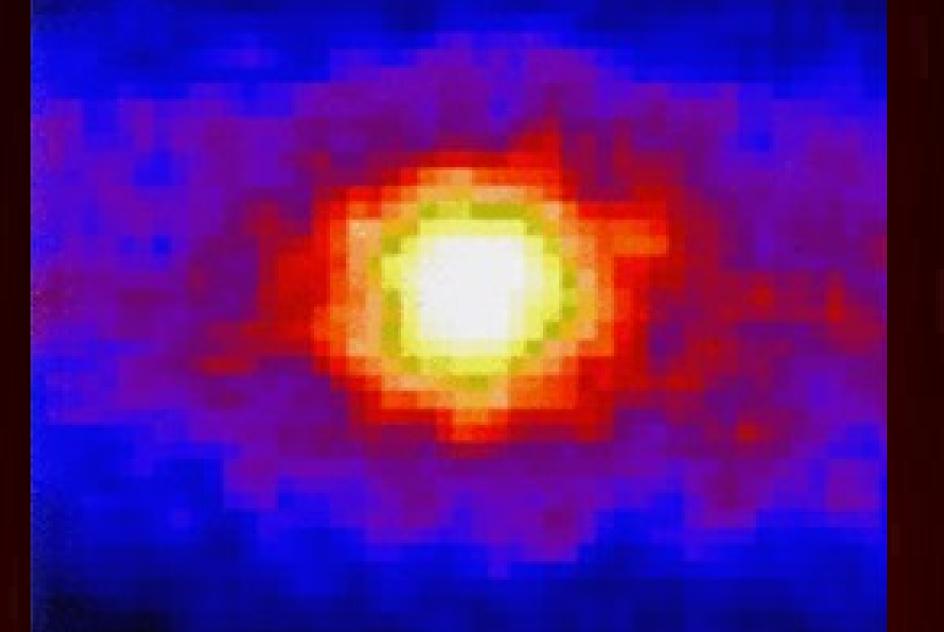
Geoneutrinos





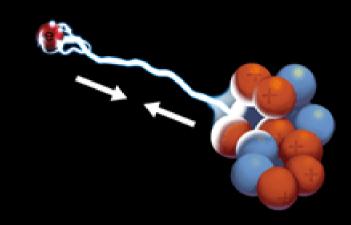


From the Sun

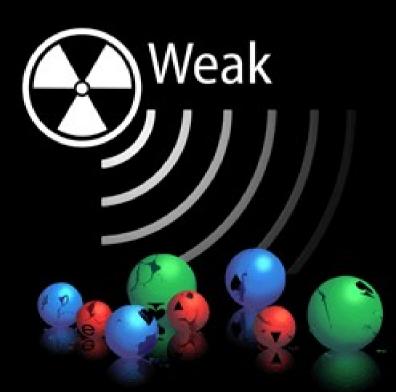








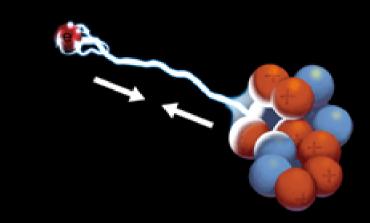


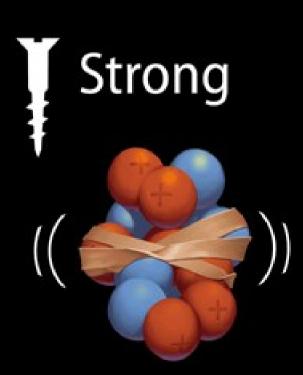


e

p V



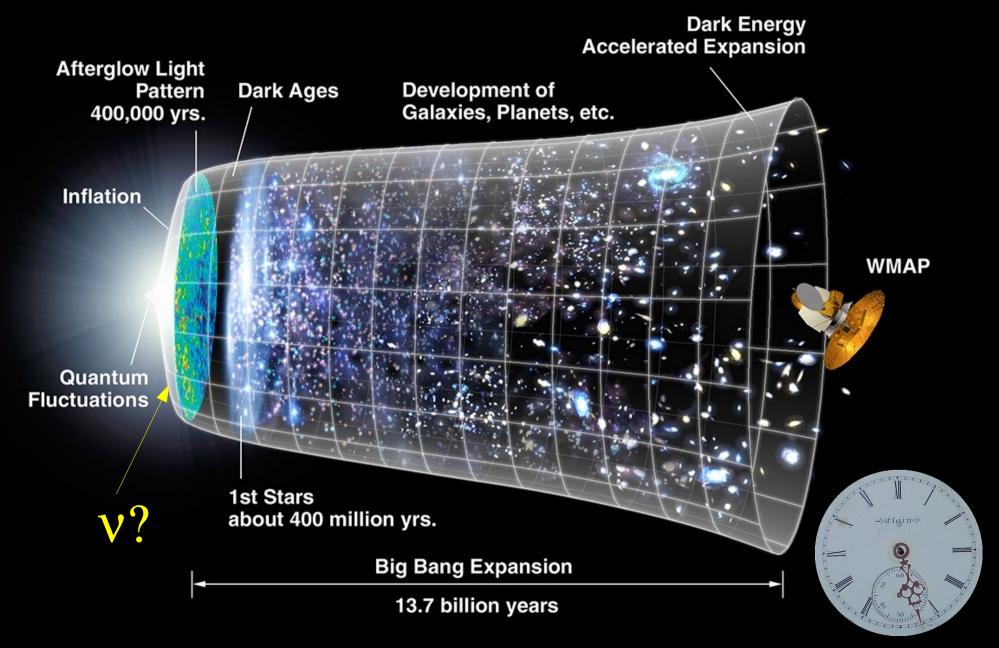


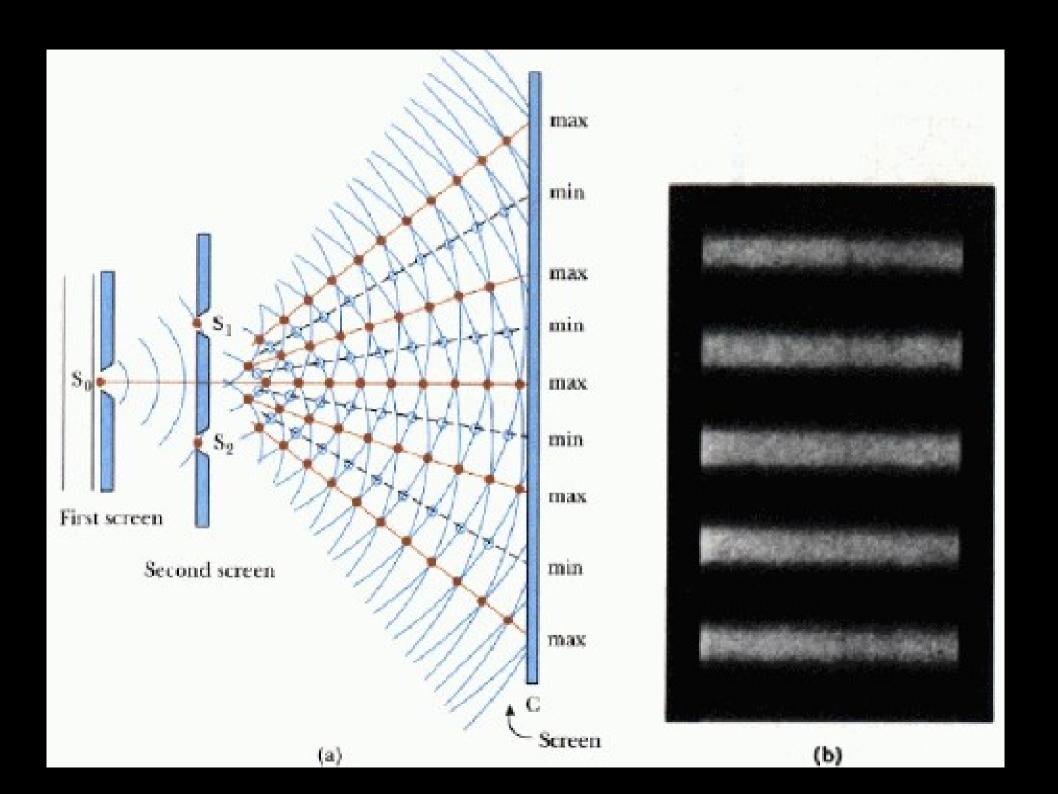


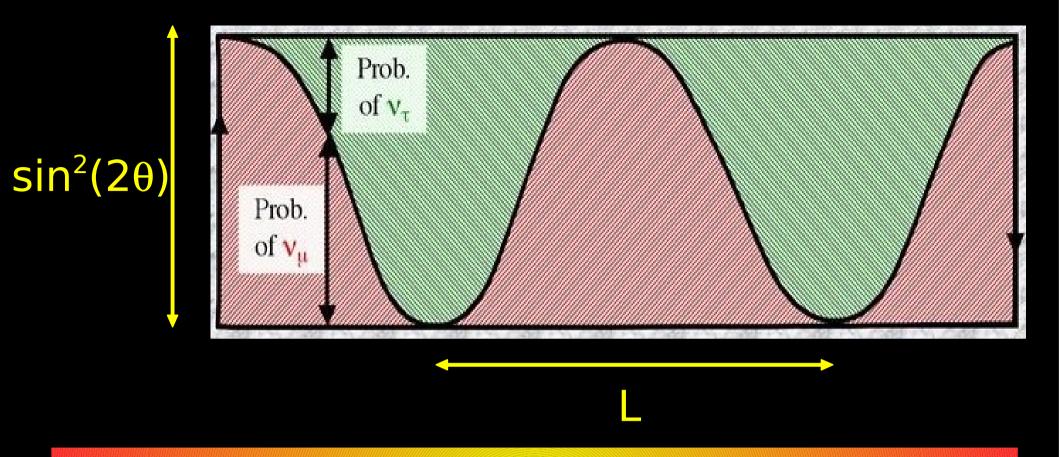


n p

I give you...the Universe







$$P(\nu_{\mu} \rightarrow \nu_{e}) = \sin^{2}(2\theta) \sin^{2}(1.27 \Delta m^{2} \frac{L}{E})$$

$$\Delta \dot{m}^2 = m_1^2 - m_2^2$$

Water Cerenkov

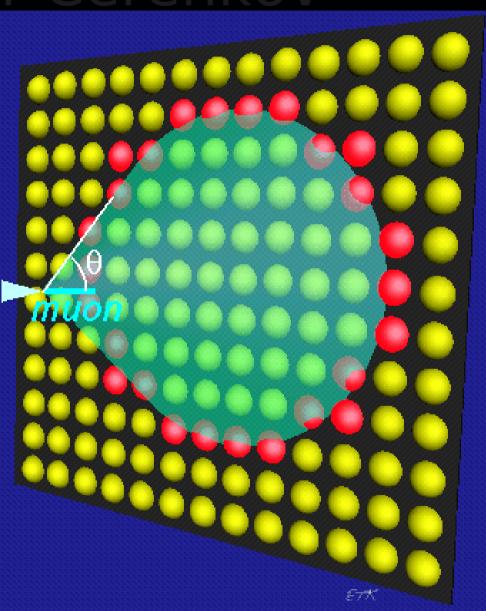
CHERENKOV EFFECT

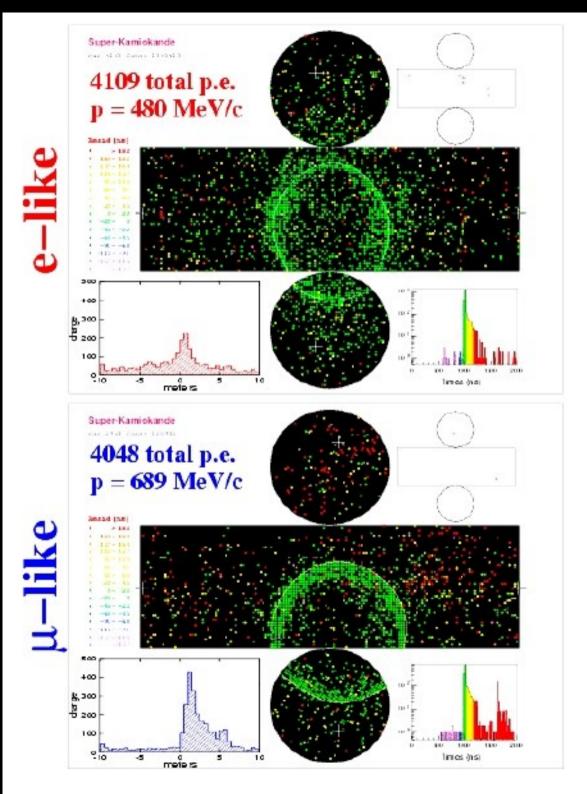
 $\beta = \mathbf{v/c}$ n(water) = 1.33

 $\cos \theta = 1/\beta n$

 $\beta = 1$ $\theta = 42$ degrees



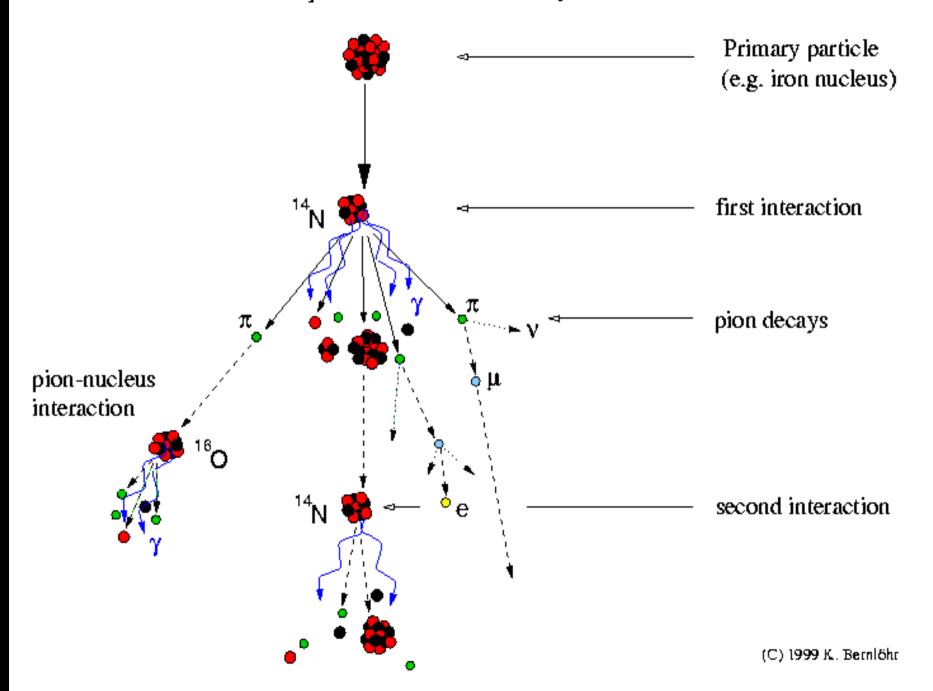


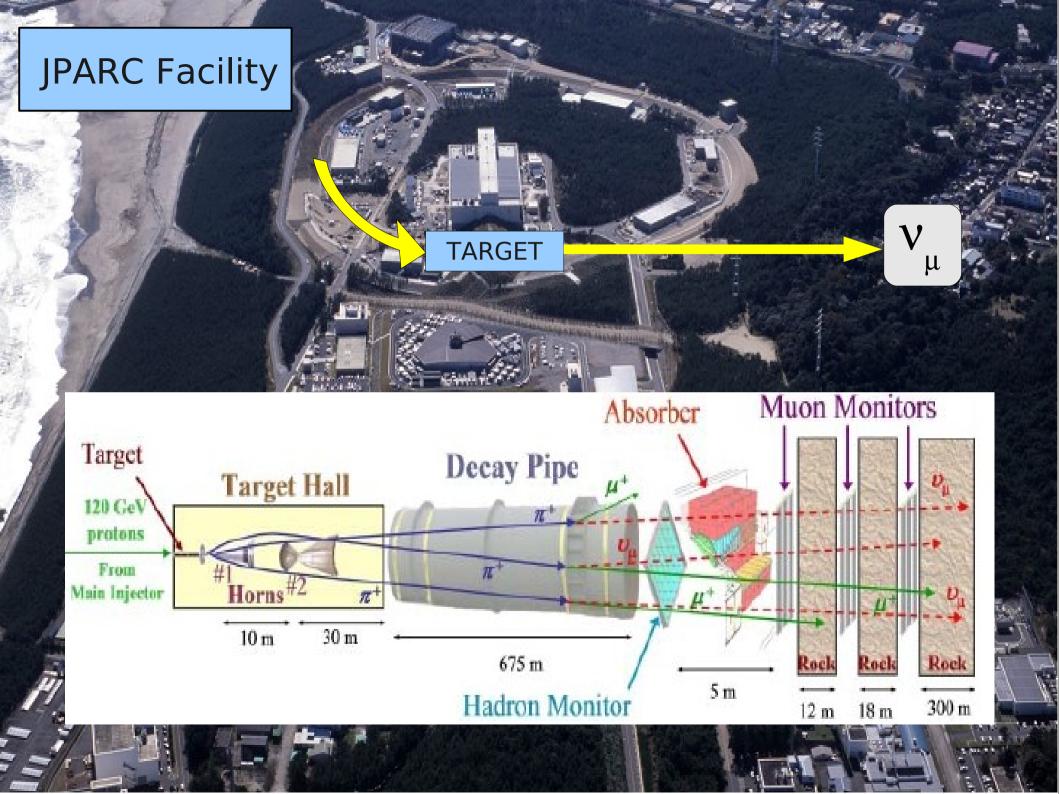


Electron-like: has a fuzzy ring

Muon-like: has a sharp edged ring and particle stopped in detector.

Development of cosmic-ray air showers





Positron, e⁺ mass (1)



Electron
Antineutrino, v

Muon, μ⁺ mass (200)



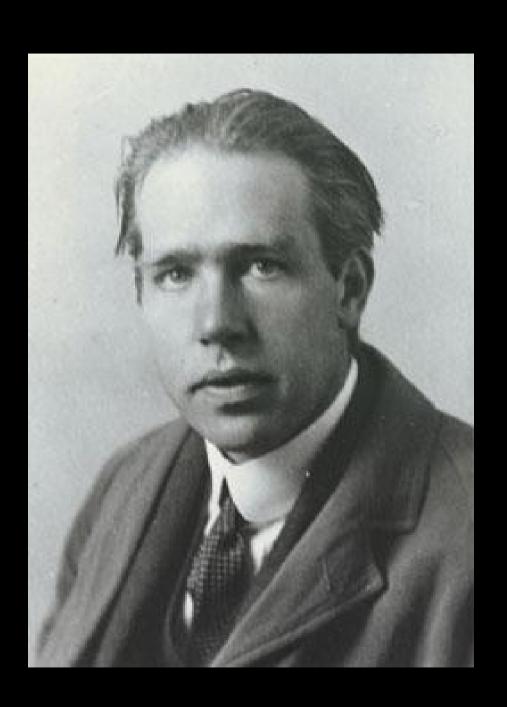
Muon Antineutrino, v

Tau, τ⁺
mass (3500)



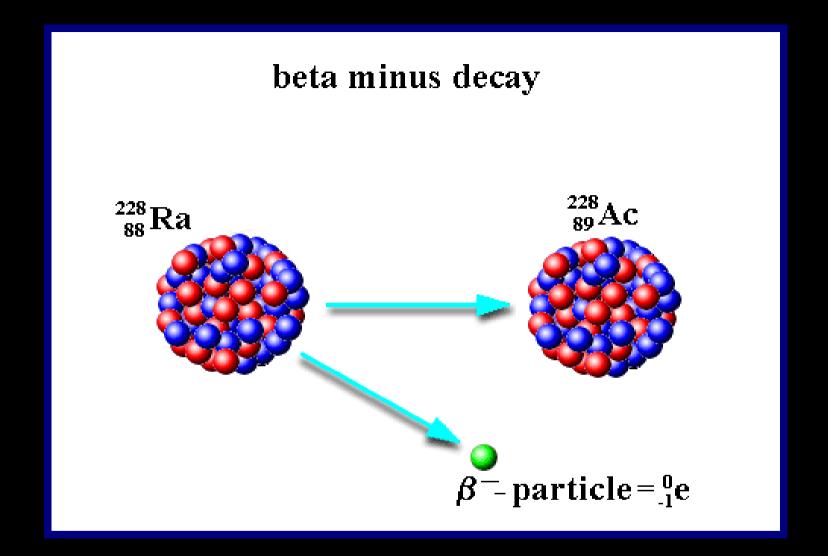
Antineutrino, v_{μ}

3 Antiparticles



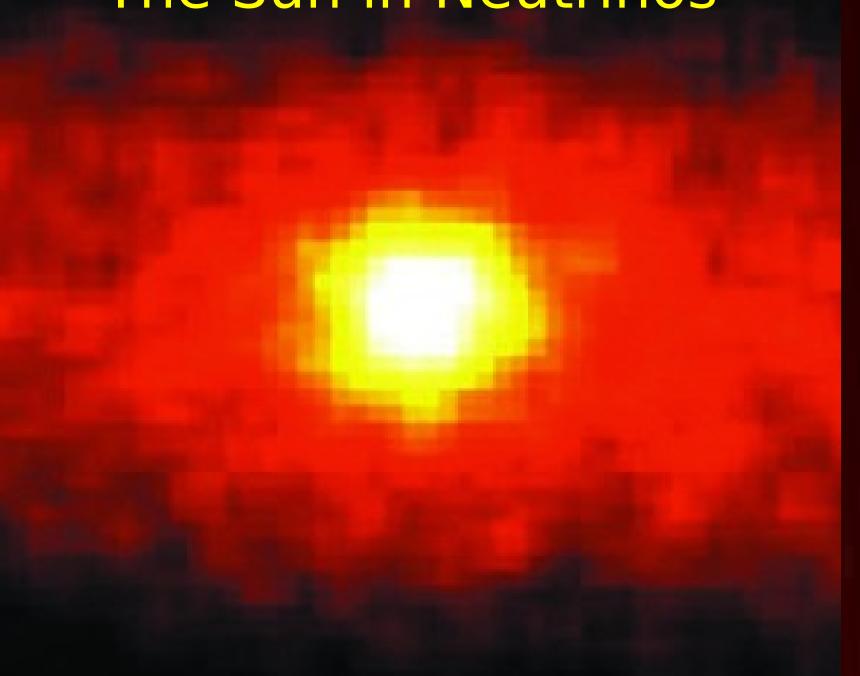
Neils Bohr

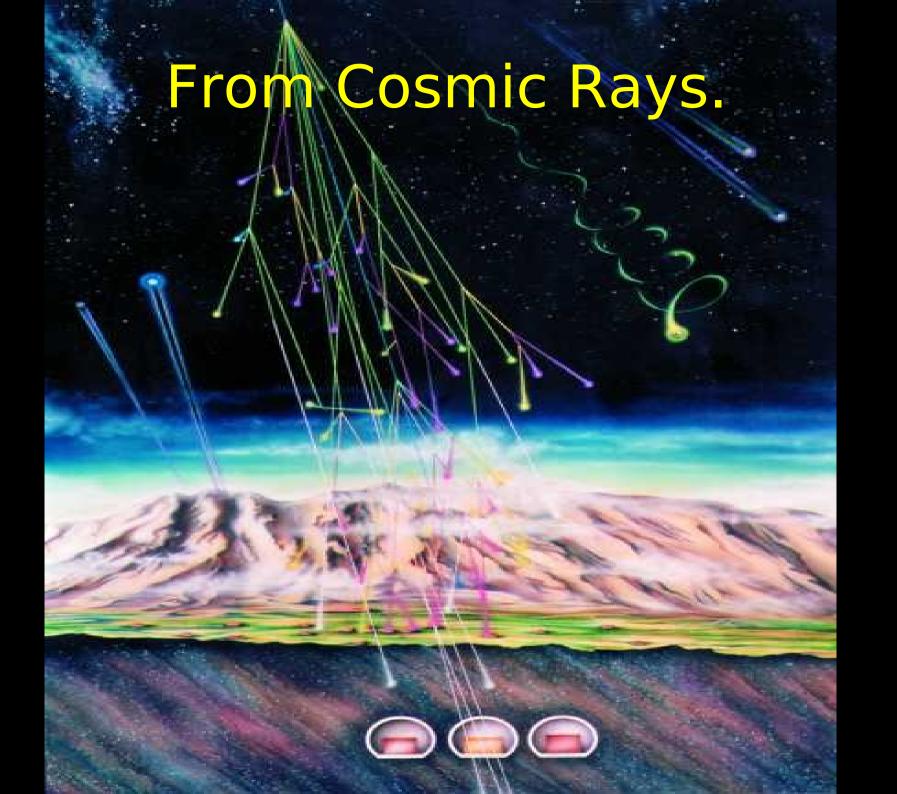
"At the present stage of atomic theory we have no arguments for upholding the concept of energy balance in the case of β -ray disintegrations."

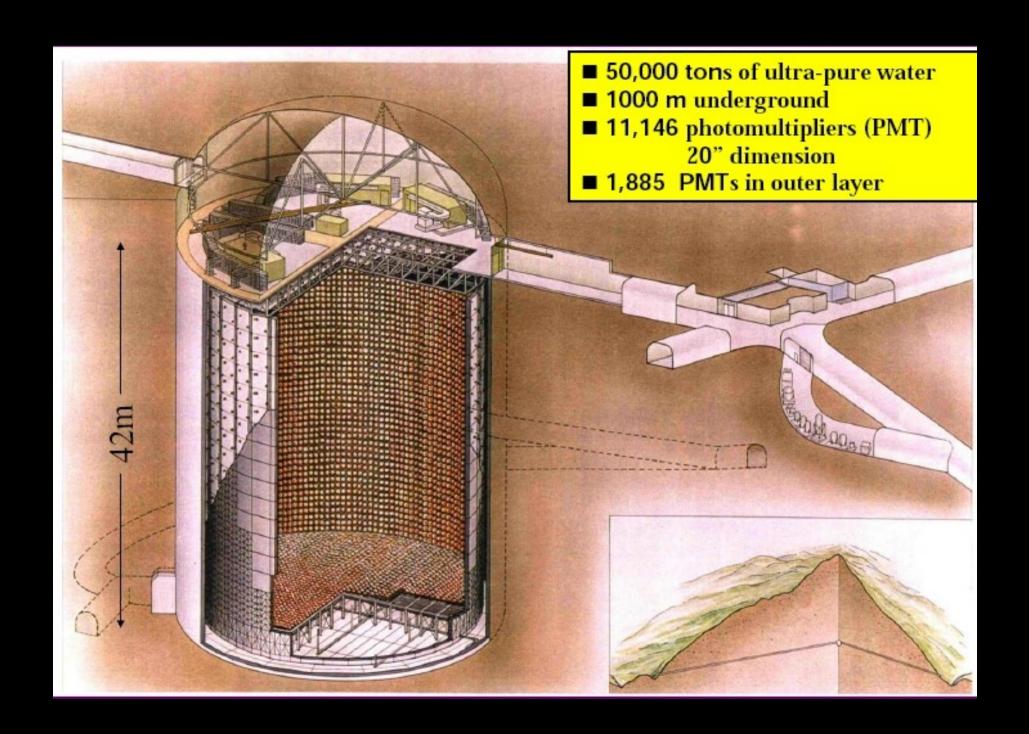


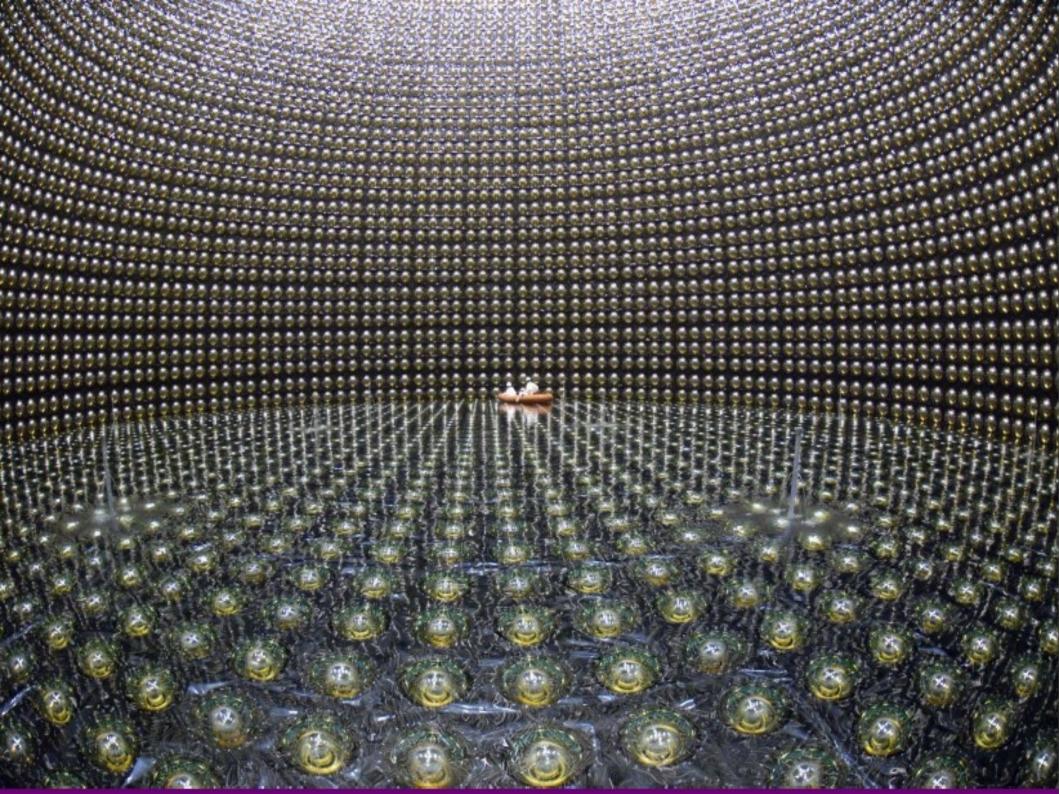
Energy(Ra) \neq Energy(Ac)+Energy(e)



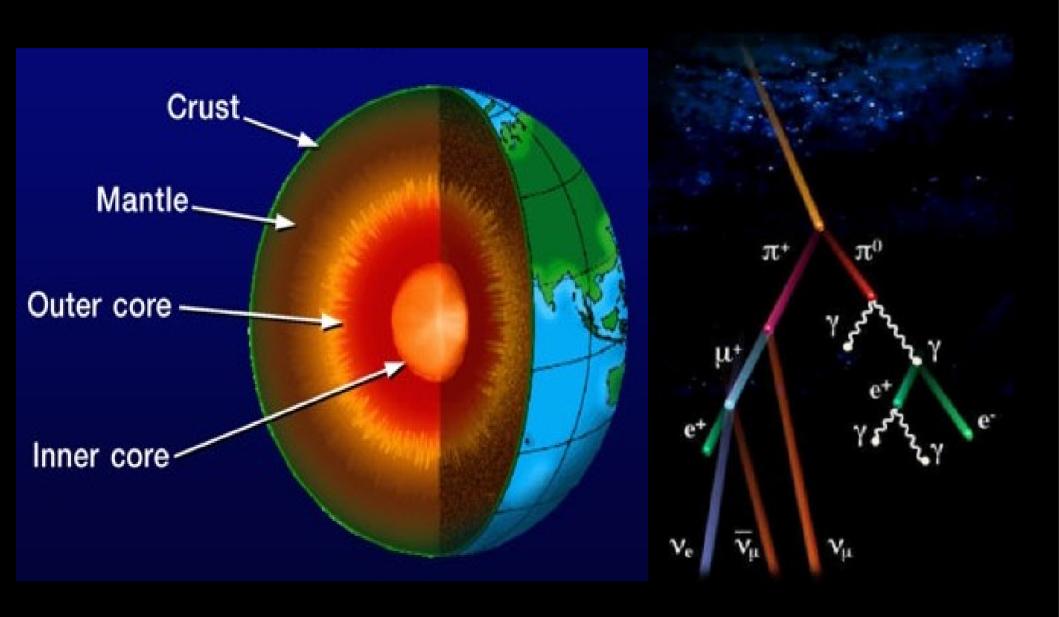






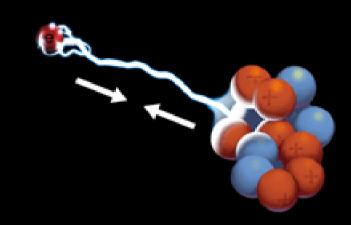


From The Earth

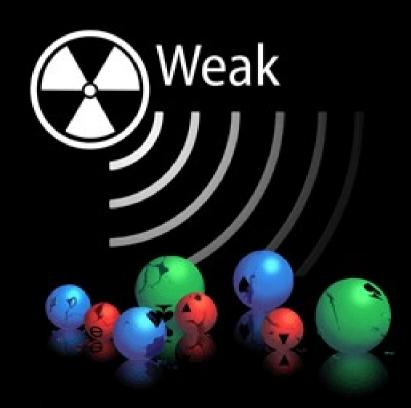












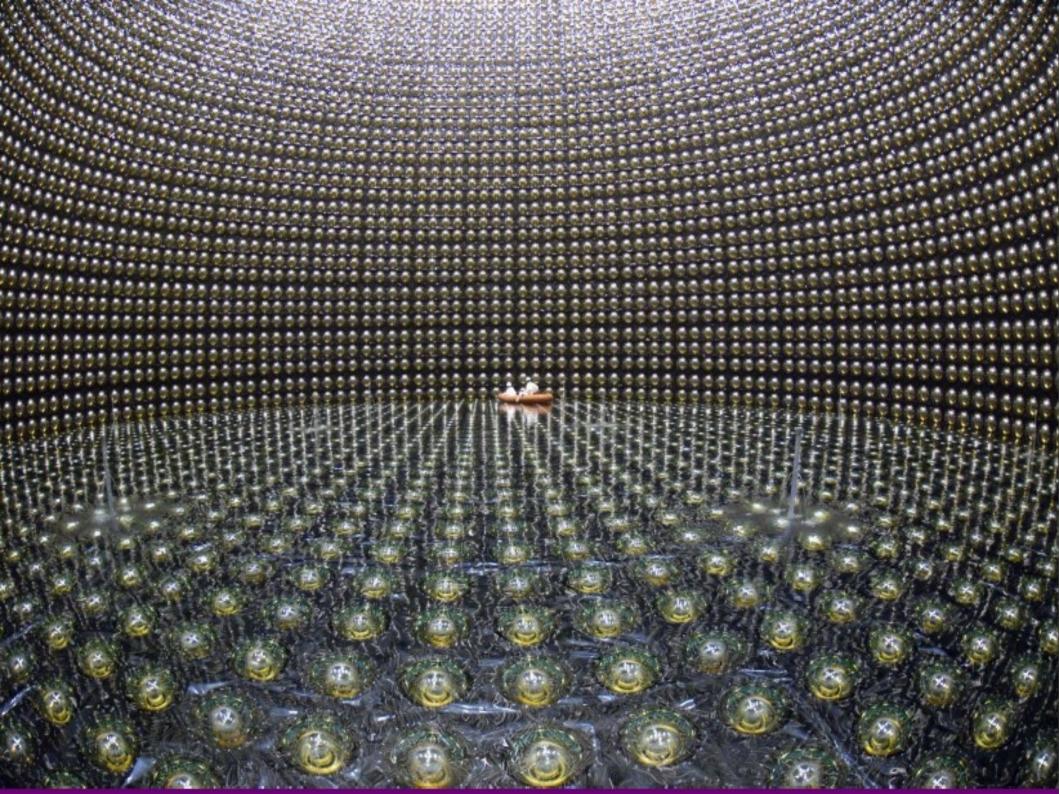
e n p

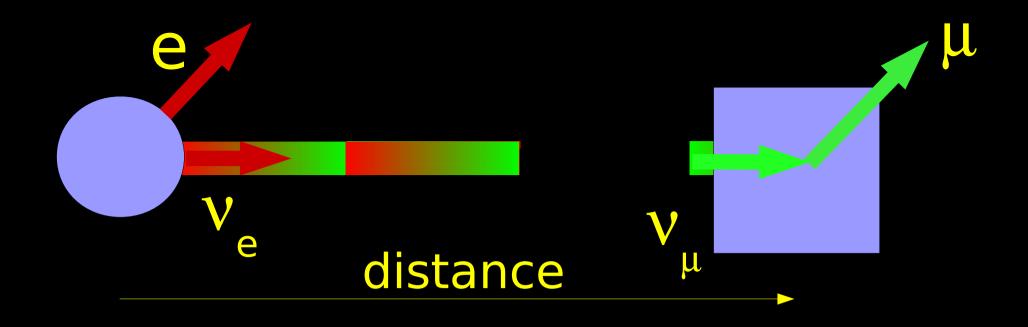


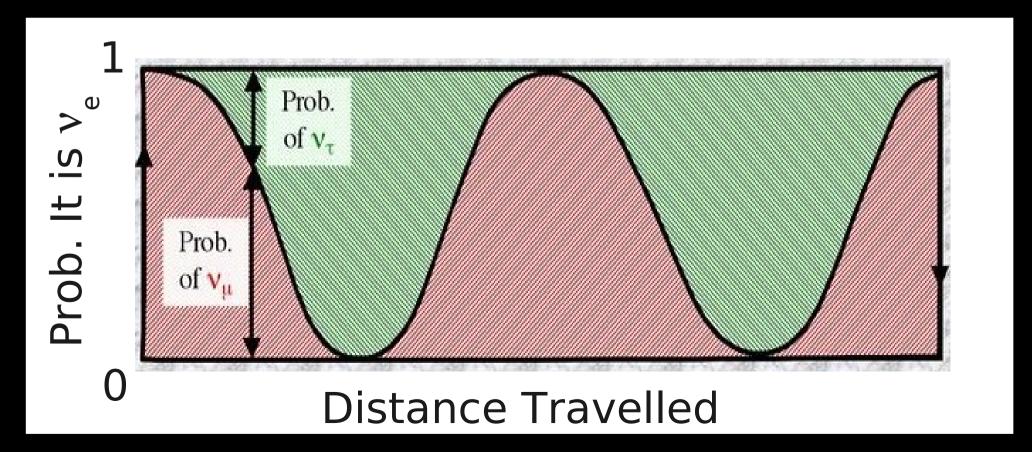


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T2K



SuperKamiokande

JPARC

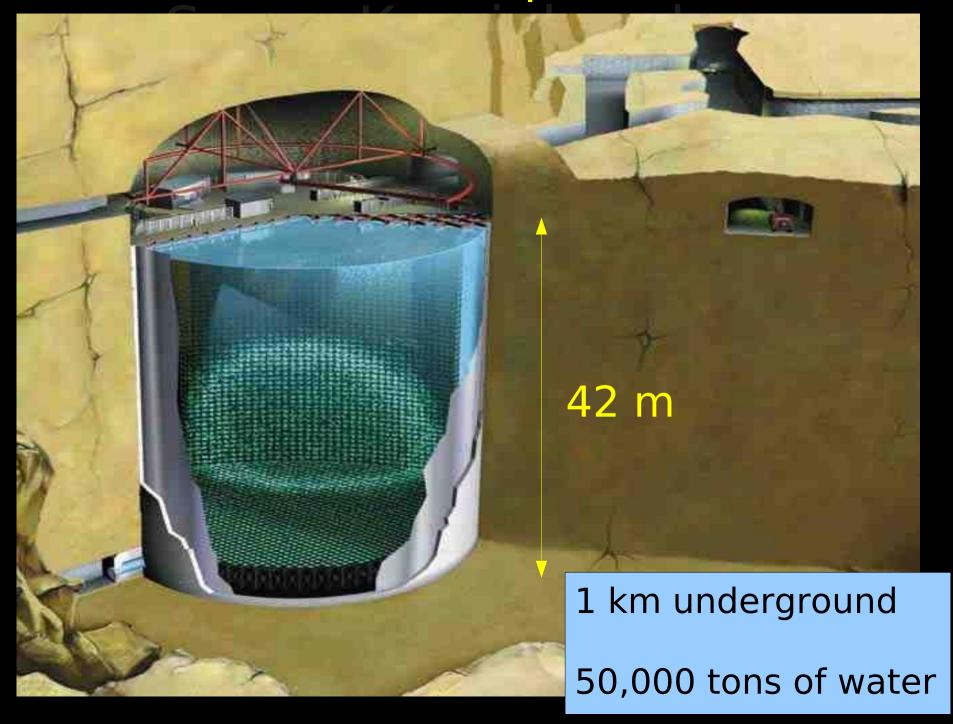
295 km

Image € 2008 TerraMetrics Image NASA

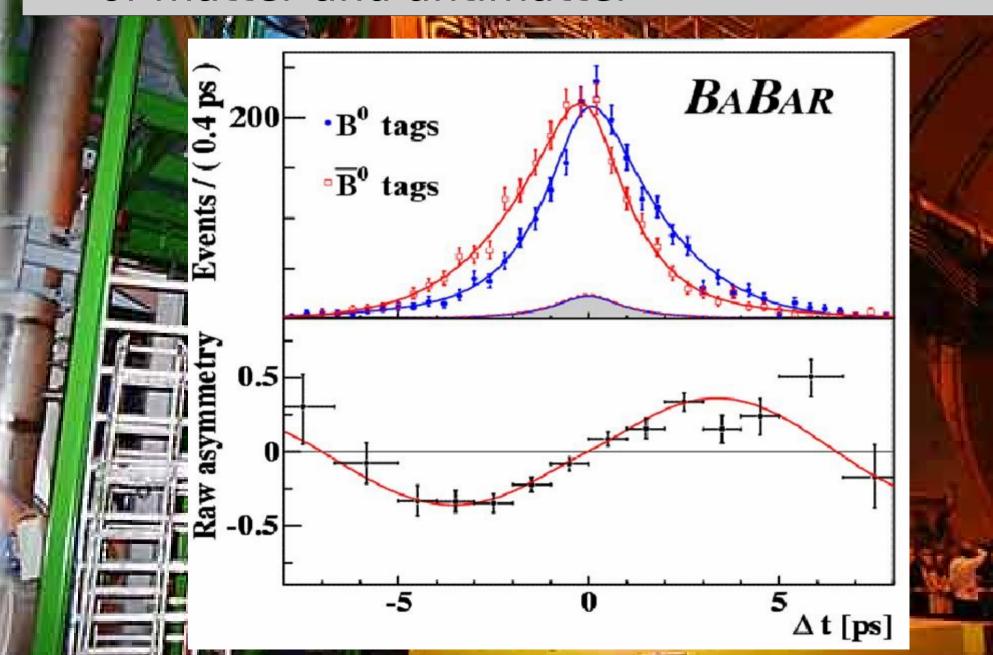
Image € 2008 Digital Earth Technology

*2007 Google™

Super-Kamiokande

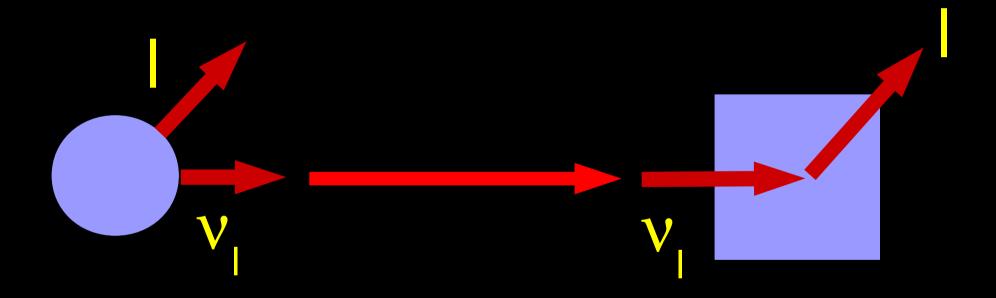


There is a difference between the physics of matter and antimatter



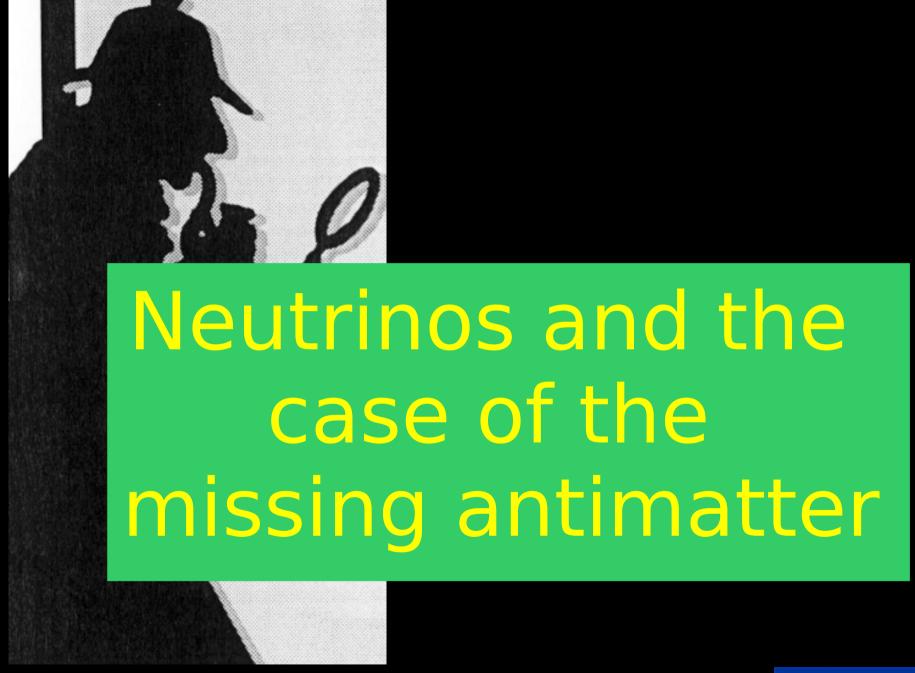
Neutrino Oscillations

THE discovery in neutrinos of the last 20 years



A typical neutrino experiment

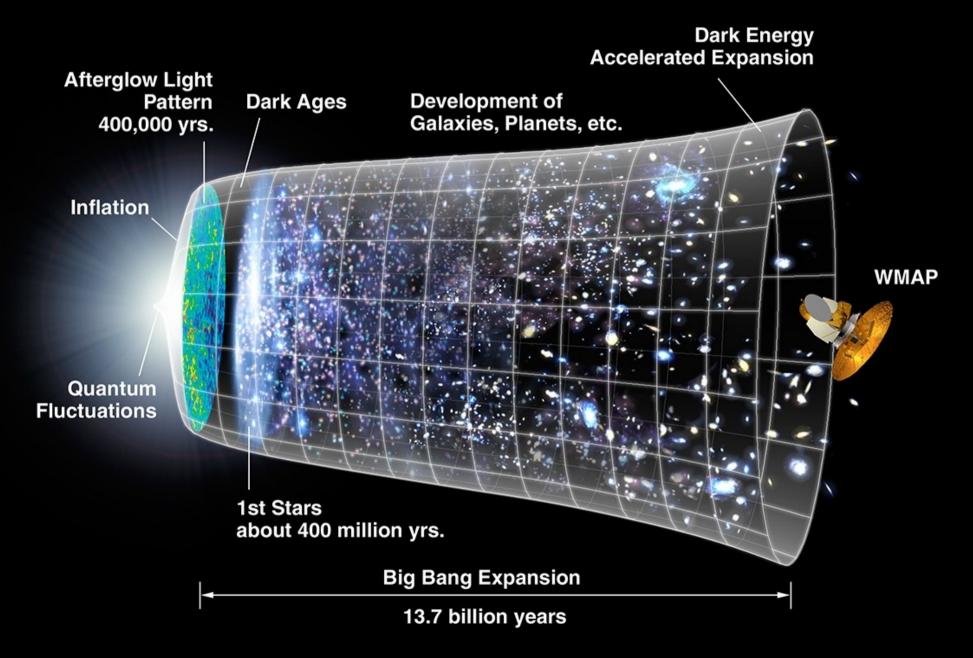






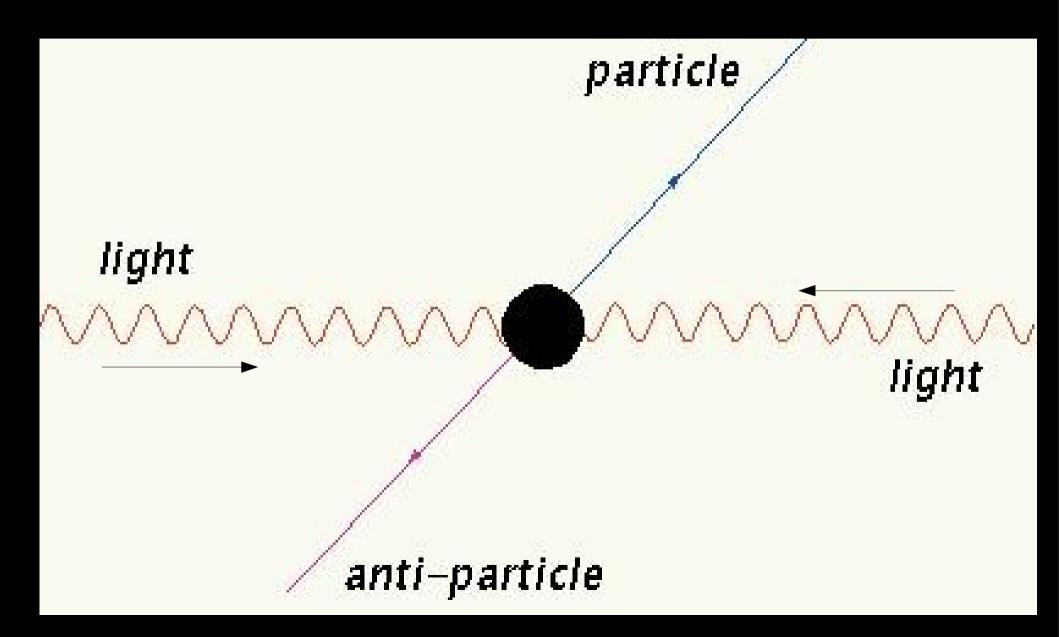
How do we exist?

The Universe

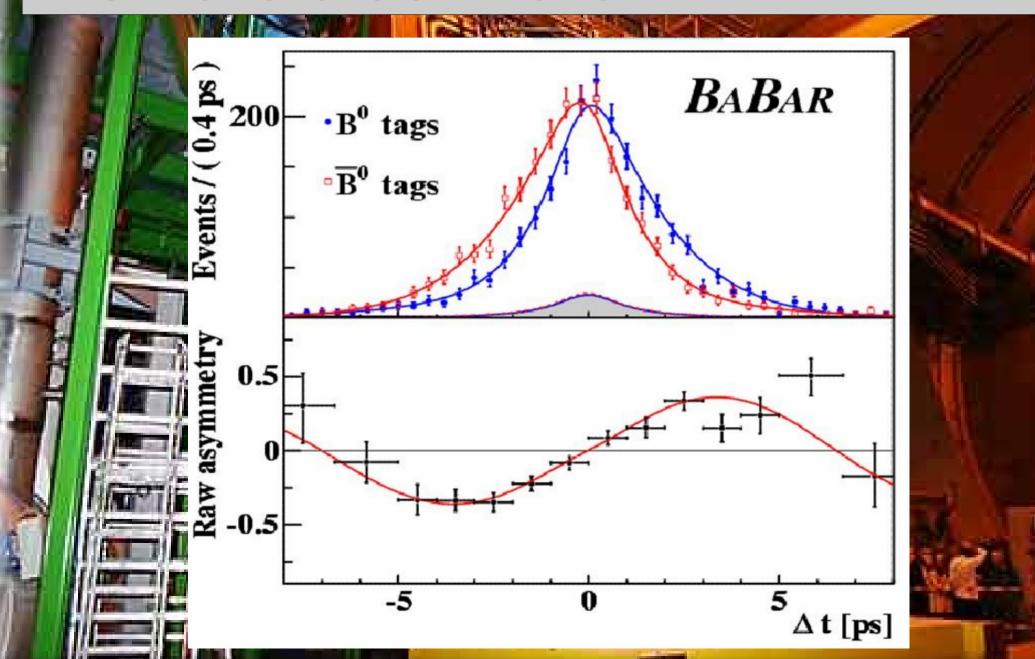


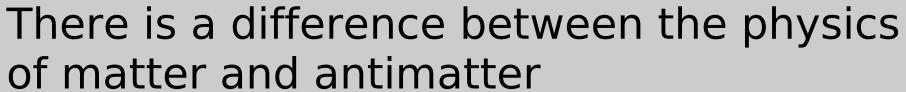


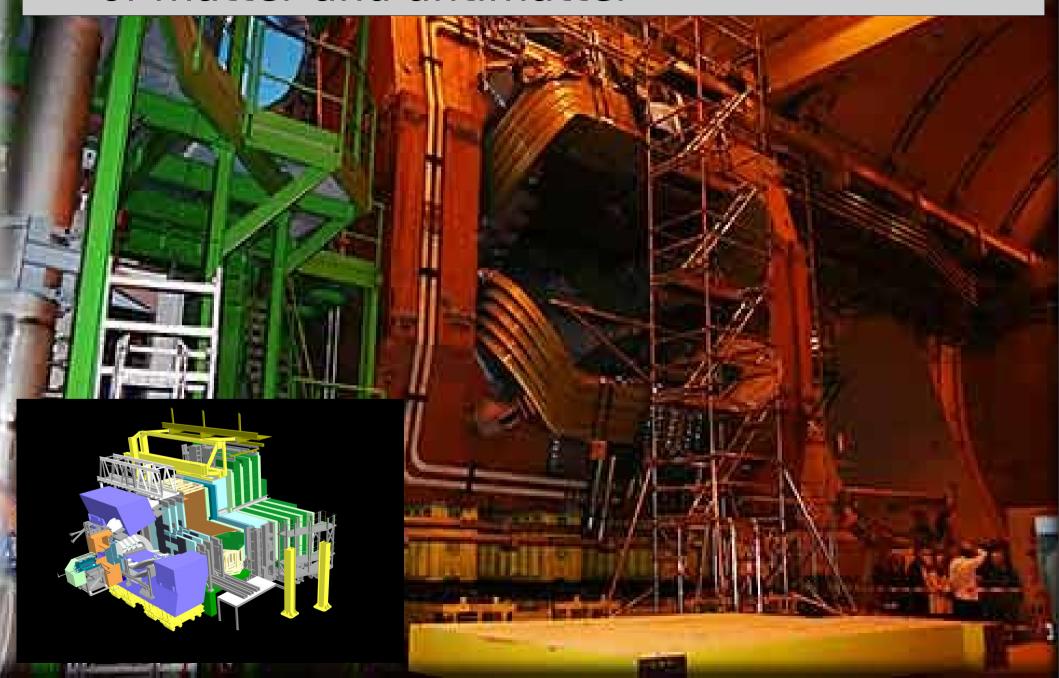
Why is there more matter than antimatter?



There is a difference between the physics of matter and antimatter







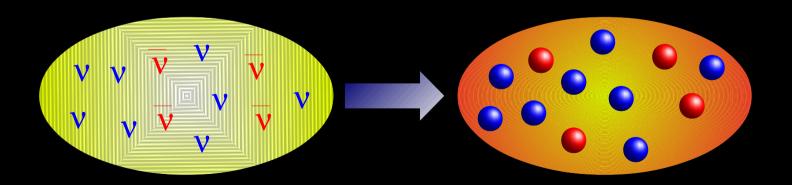
How do we exist?

We don't know (yet) but we're working on it

The smallest, most insignificant (yet most common) particle in the cosmos may just hold the reason!

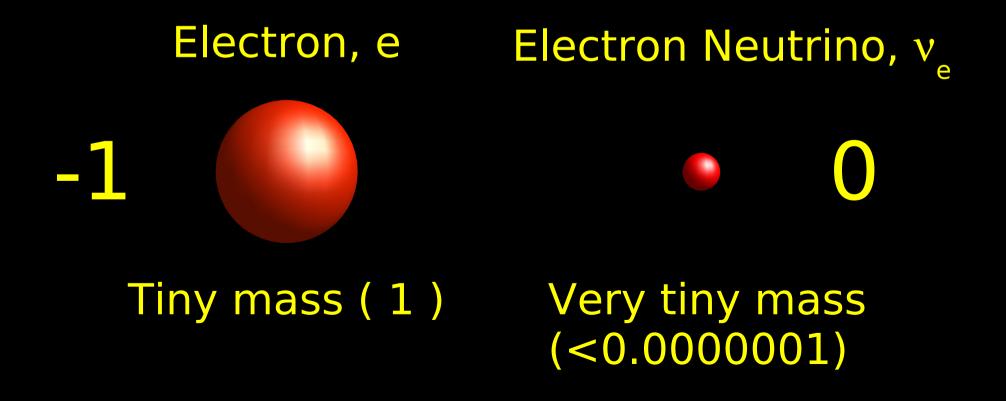
Matter-Antimatter Asymmetry

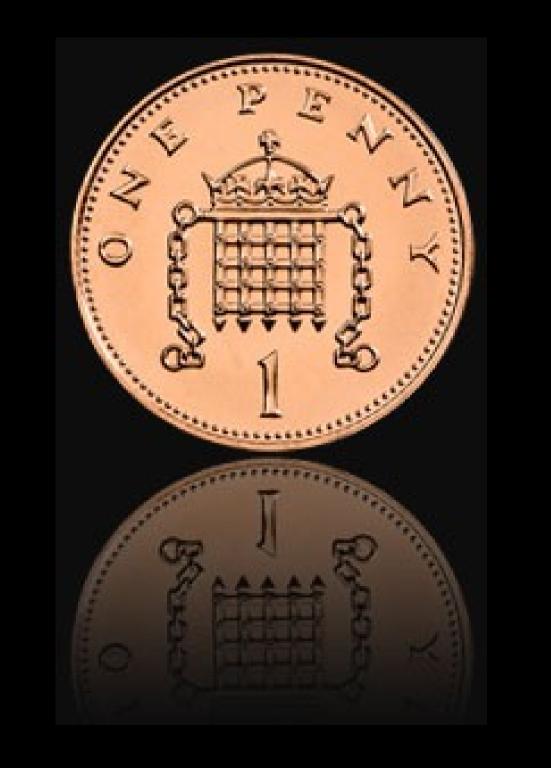
A theory called "Leptogenesis" suggests that the asymmetry we see was generated by an asymmetry between *neutrinos* and *anti-neutrinos* at the beginning of things.



What is a neutrino?

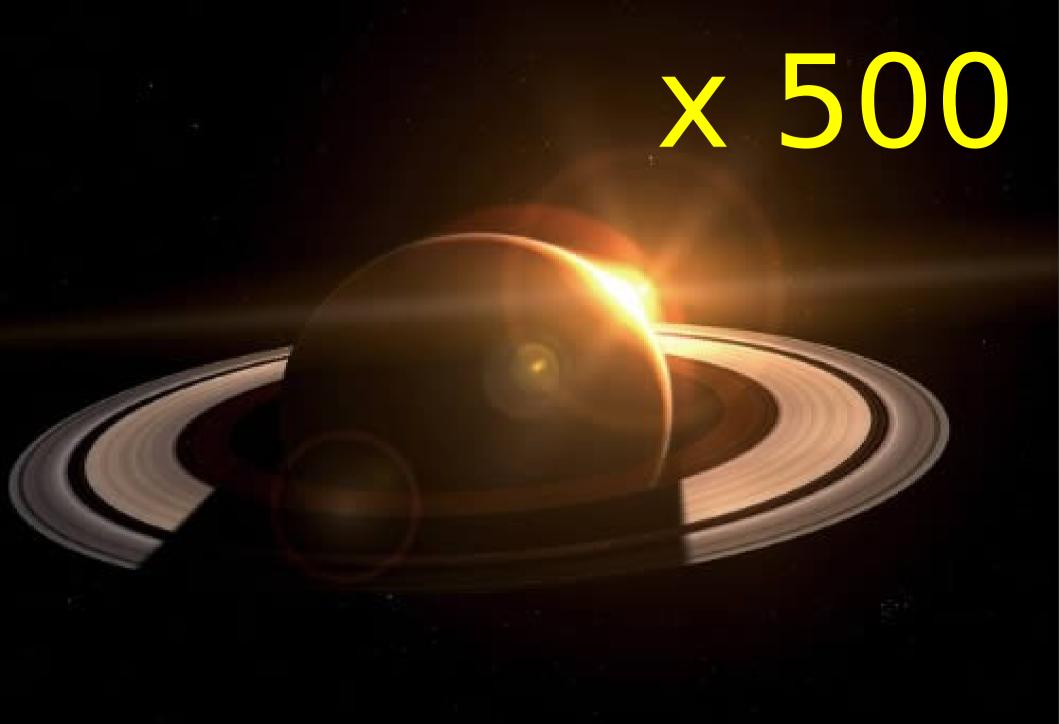
Neutrinos are the second most common particle in the universe. Produced whereever you have radioactive decays

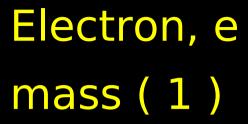














Electron Neutrino, v_e

Muon, μ mass (200)



Muon Neutrino, ν_μ

Tau, τ mass (3500)



Tau Neutrino, $v_{_{\tau}}$

3 Lepton Types

Electron Neutrino, v_e



Electron
Antineutrino, \overline{v}_e

Muon Neutrino, ν_{μ}



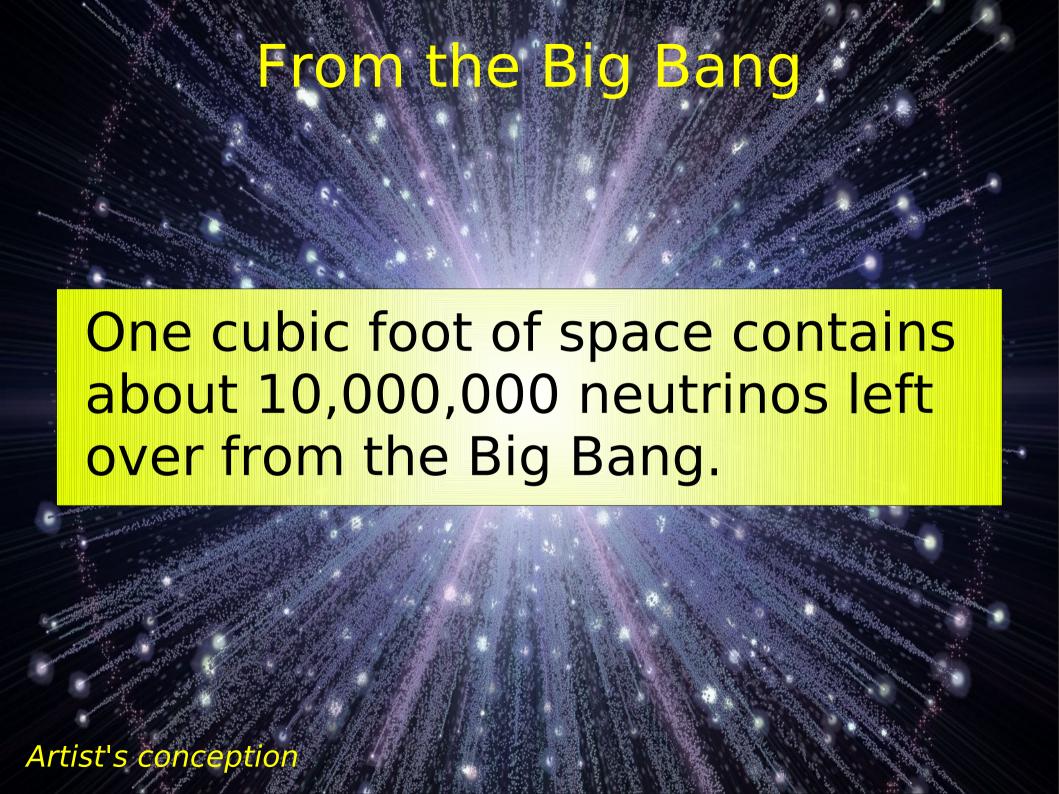
Muon
Antineutrino, $\overline{\nu}_{\mu}$

Tau Neutrino, v_{τ}



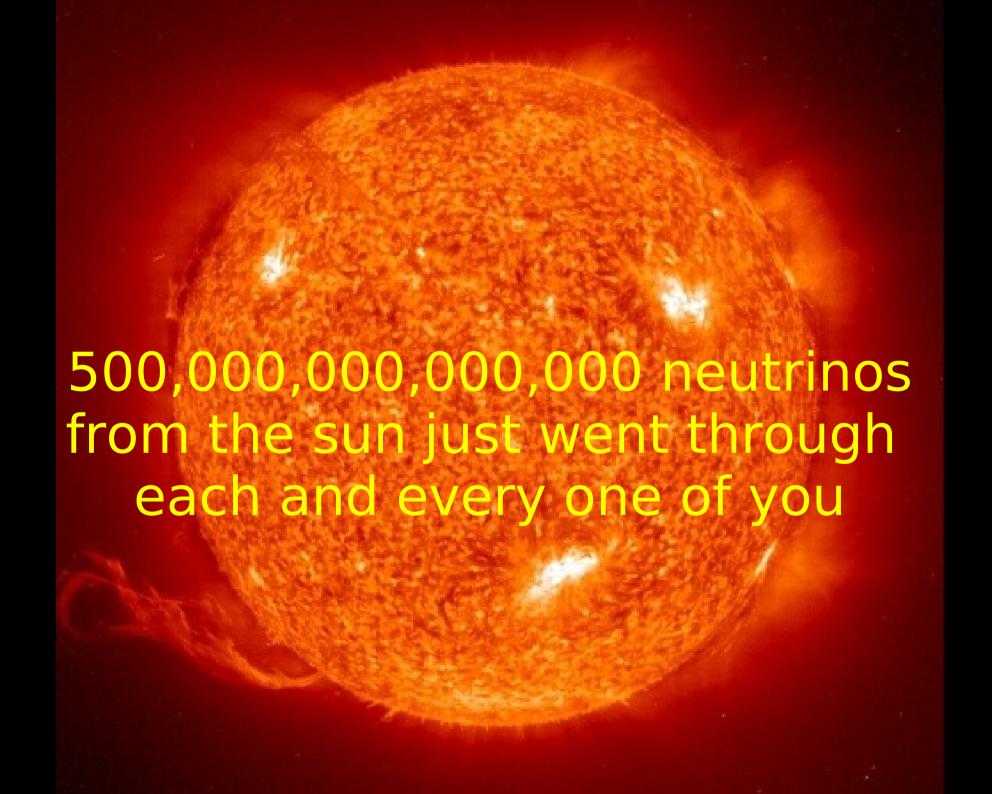
Antineutrino, $\overline{\nu}_{\mu}$

3 neutrino Flavours

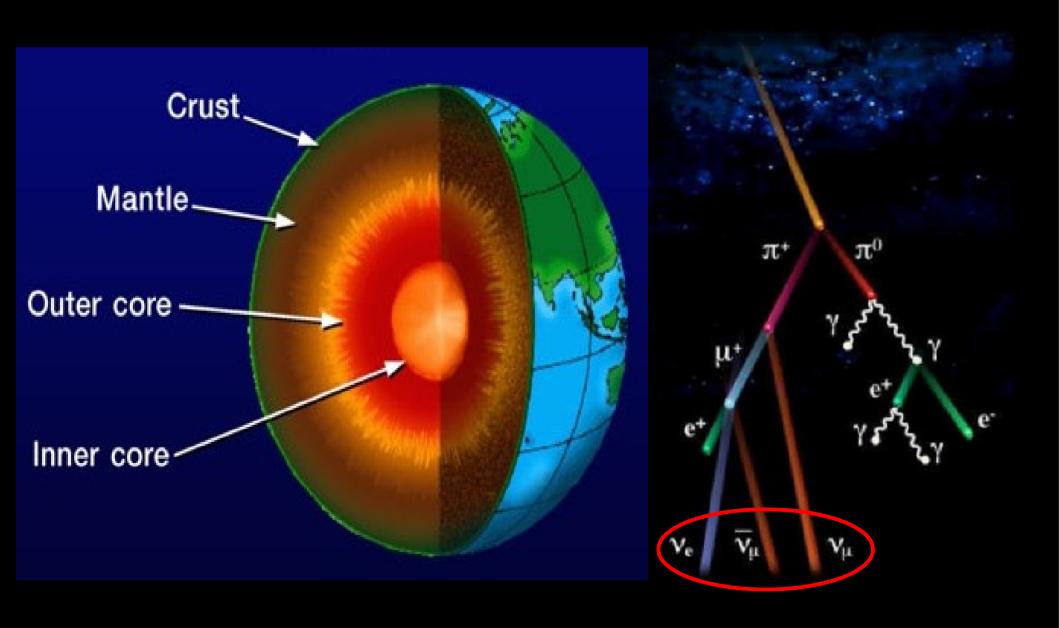




Supernovae created the heavy elements (us) and neutrinos may be responsible for the star exploding.



From The Earth





So why don't we notice?

v are almost ghosts. They interact extremely weakly with matter.

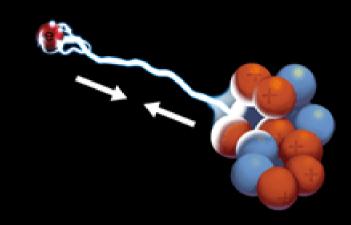
To a neutrino a planet is mostly empty space.

"The chances of a neutrino actually hitting something as it travels through all this howling emptiness are roughly comparable to that of dropping a ball bearing at random from a cruising 747 and hitting, say, an egg sandwich."

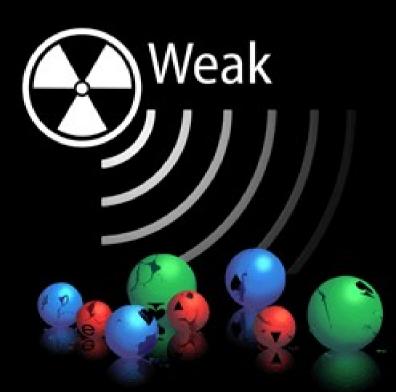
Douglas Adams











e

p V Assume 1 billion people eat an egg sandwich every 3 months

1.67 x 10⁷ egg sandwiches/day

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 1.67×10^7 egg sandwiches/day

Let's say that 3 months of the year people can eat outside, and that they picnic one day every week 600,000 external egg sandwiches/day

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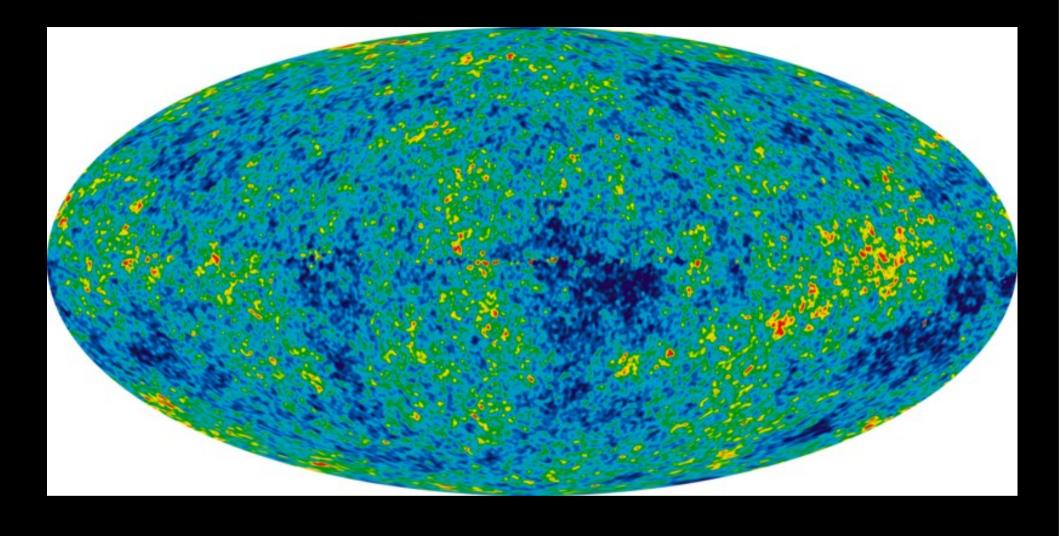
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| Area of egg sandwich – 15 cm x 15cm | 186 m² total egg-sandwich area |
| Surface area of earth | 500 million km ² |
| Suppose flight paths cover area of earth uniformly | Probability of egg-sandwich/ball bearing intersection 3 x 10 ⁻¹³ |

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intersection

Probability of average solar neutrino interaction

 5×10^{-13}



One cubic foot of space contains about 10,000,000 neutrinos left over from the Big Bang.