

# Neutrinos and the Case of the Missing Antimatter

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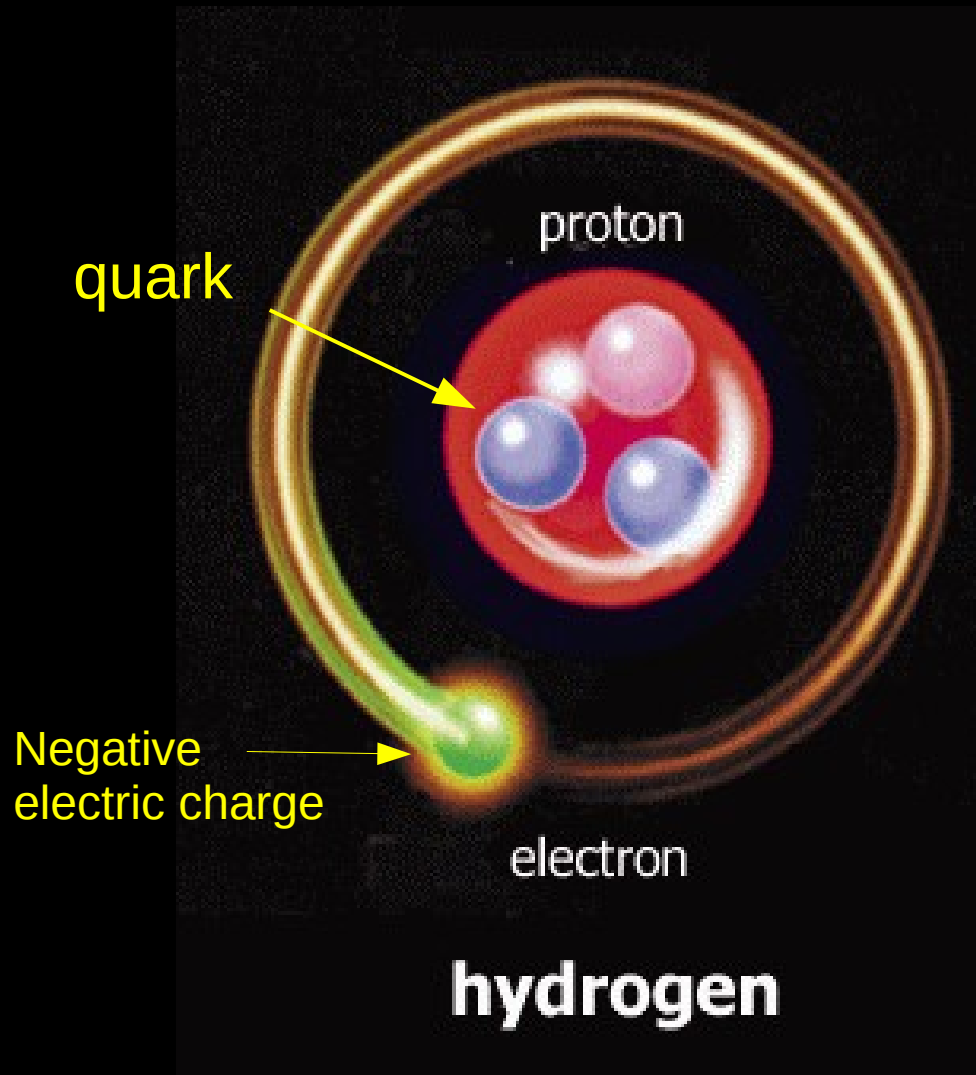




A vast field of stars, primarily blue and white, scattered across a dark cosmic background. The stars vary in brightness and size, creating a rich, textured appearance. The text 'Where is all the antimatter?' is centered horizontally and rendered in a bright yellow, sans-serif font.

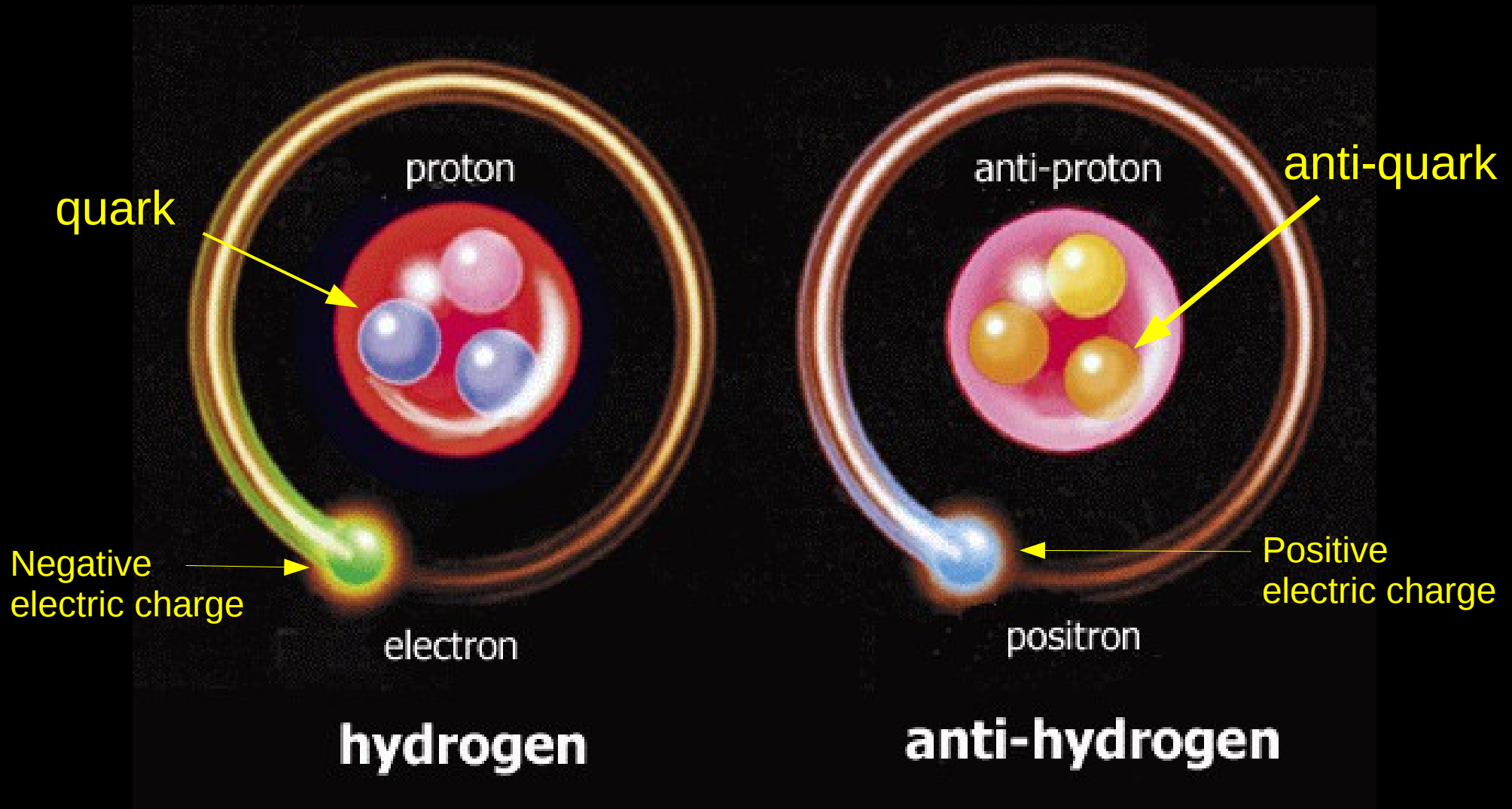
Where is all the antimatter?

# Matter



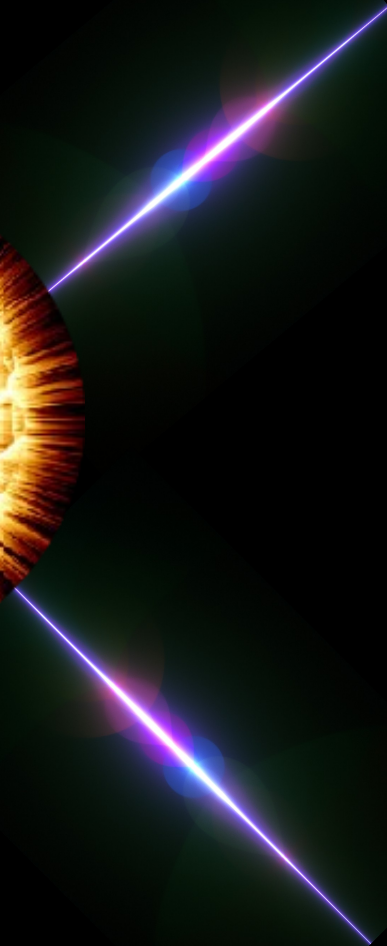
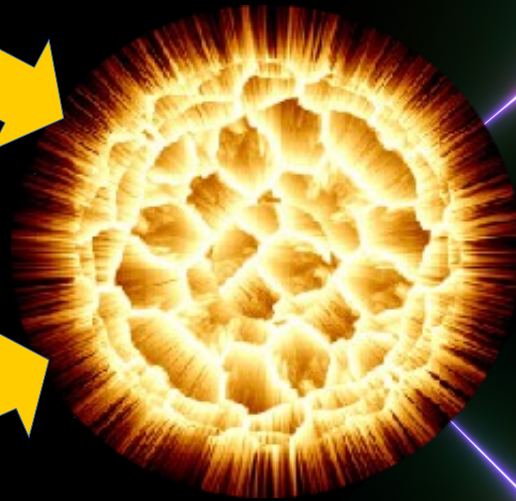
# Matter

# Anti-Matter



Predicted by Dirac in 1928  
Discovered by Anderson in 1932

Matter

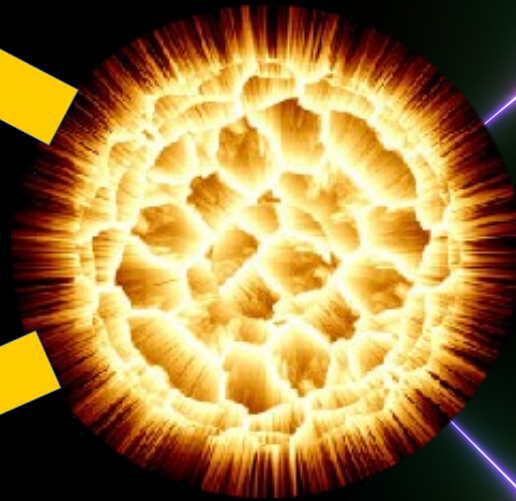


light

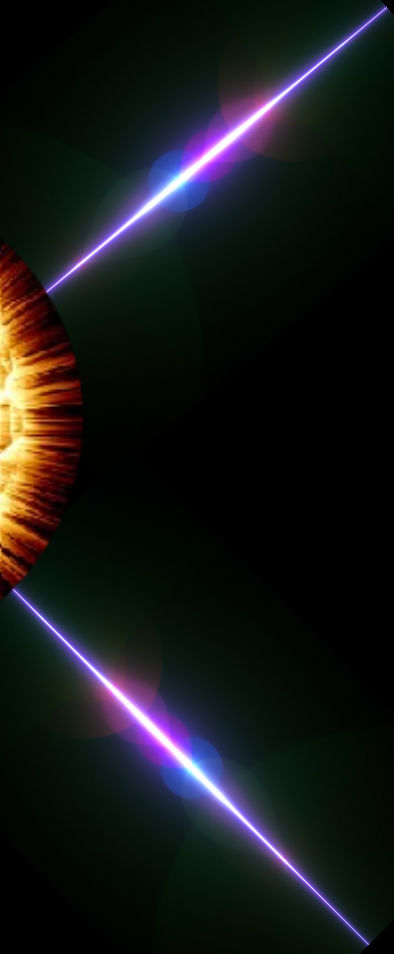
Anti-Matter

$$E = mc^2$$

Matter



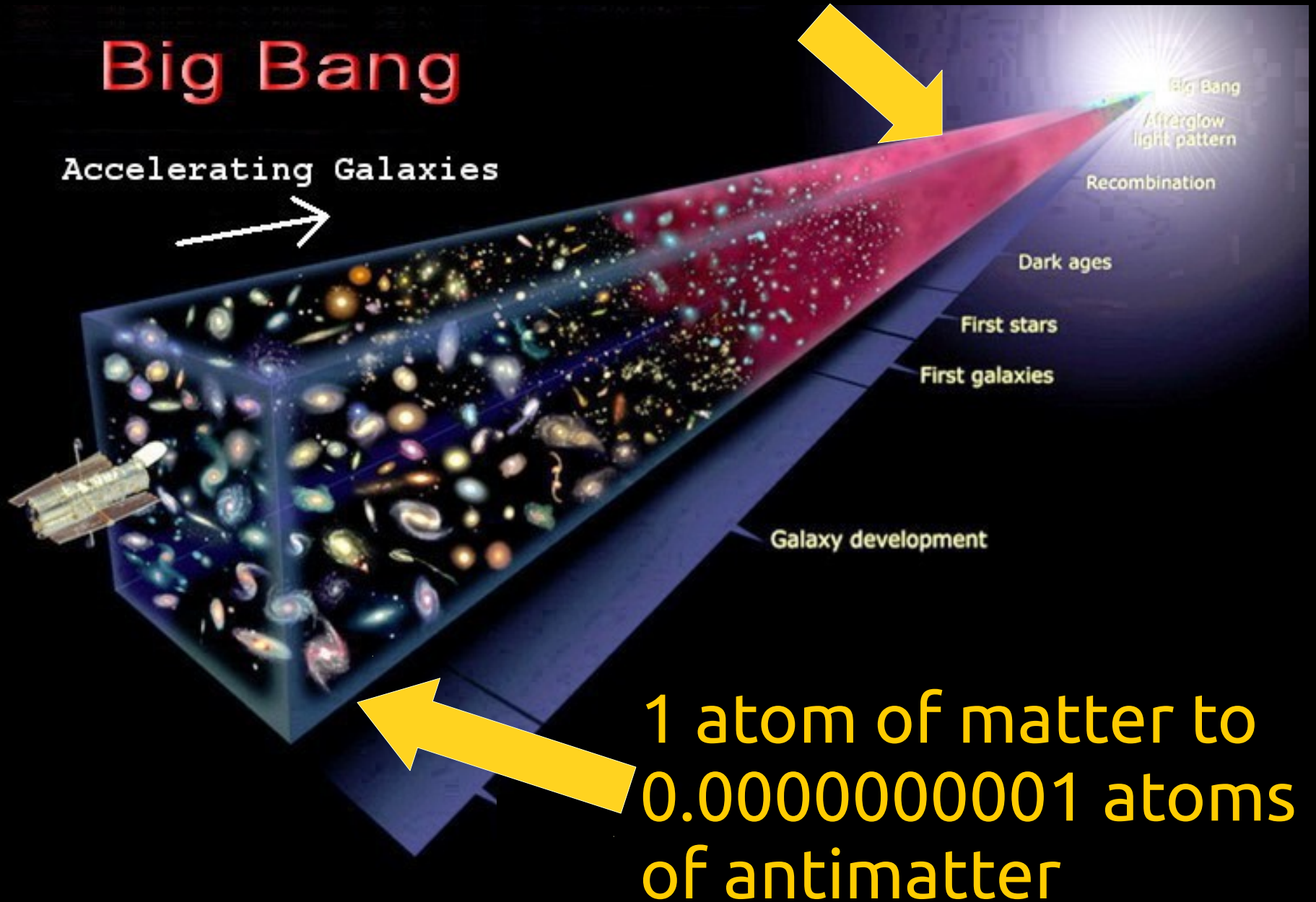
light



Anti-Matter

The reverse reaction should also happen with the same probability

Equal amounts of matter and antimatter





# Neutrinos

The smallest, most insignificant (yet most common) particle in the cosmos may just hold the reason!

This particle is called the *neutrino*

# So what is a neutrino?

Neutrinos are the second most common particle in the universe. They are produced whenever something radioactively decays

Electron,  $e$

Electron Neutrino,  $\nu_e$

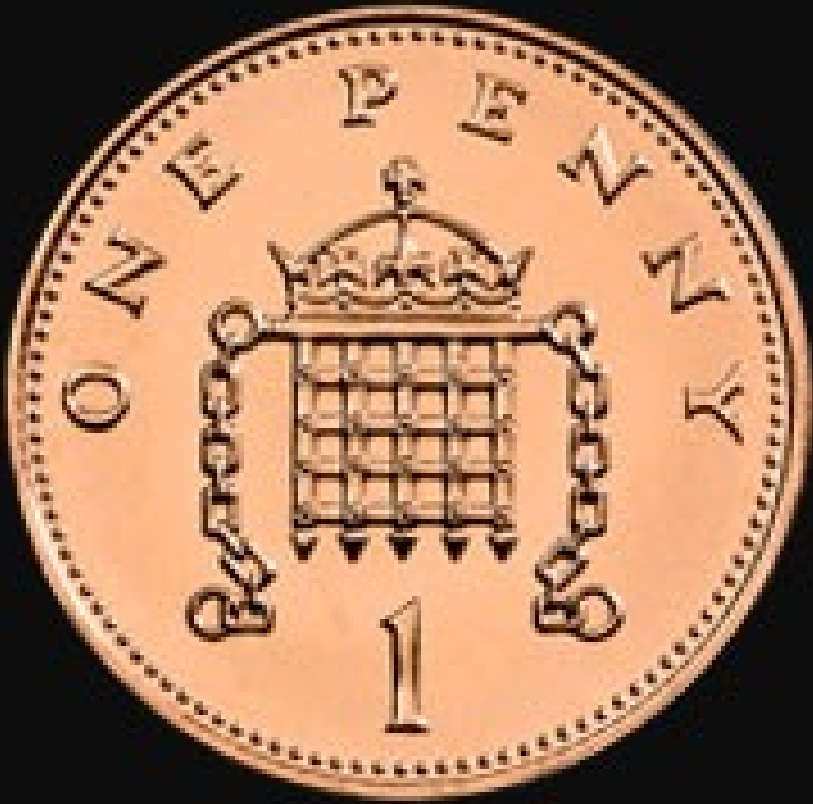
Negative  
Charge



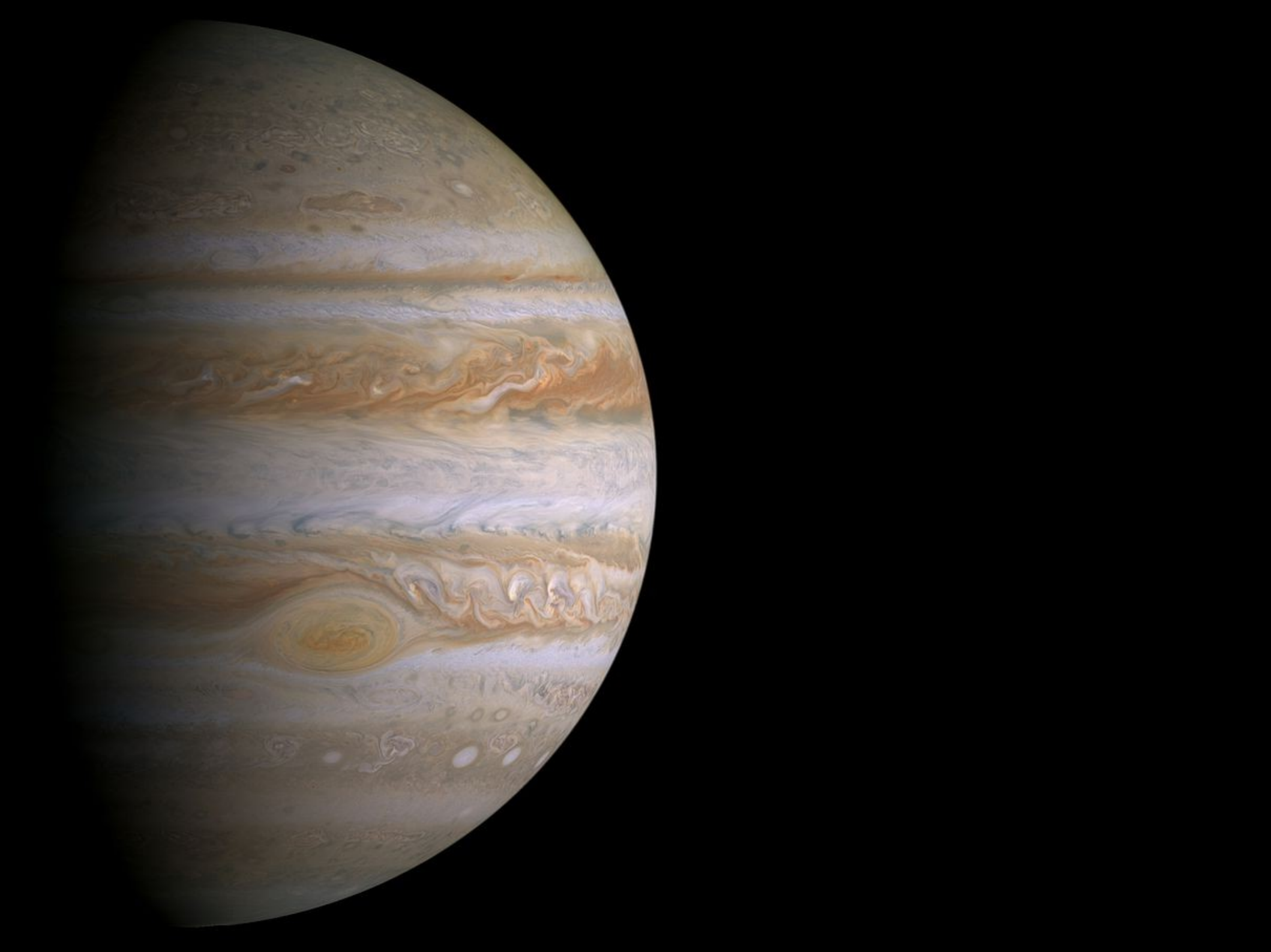
Neutral

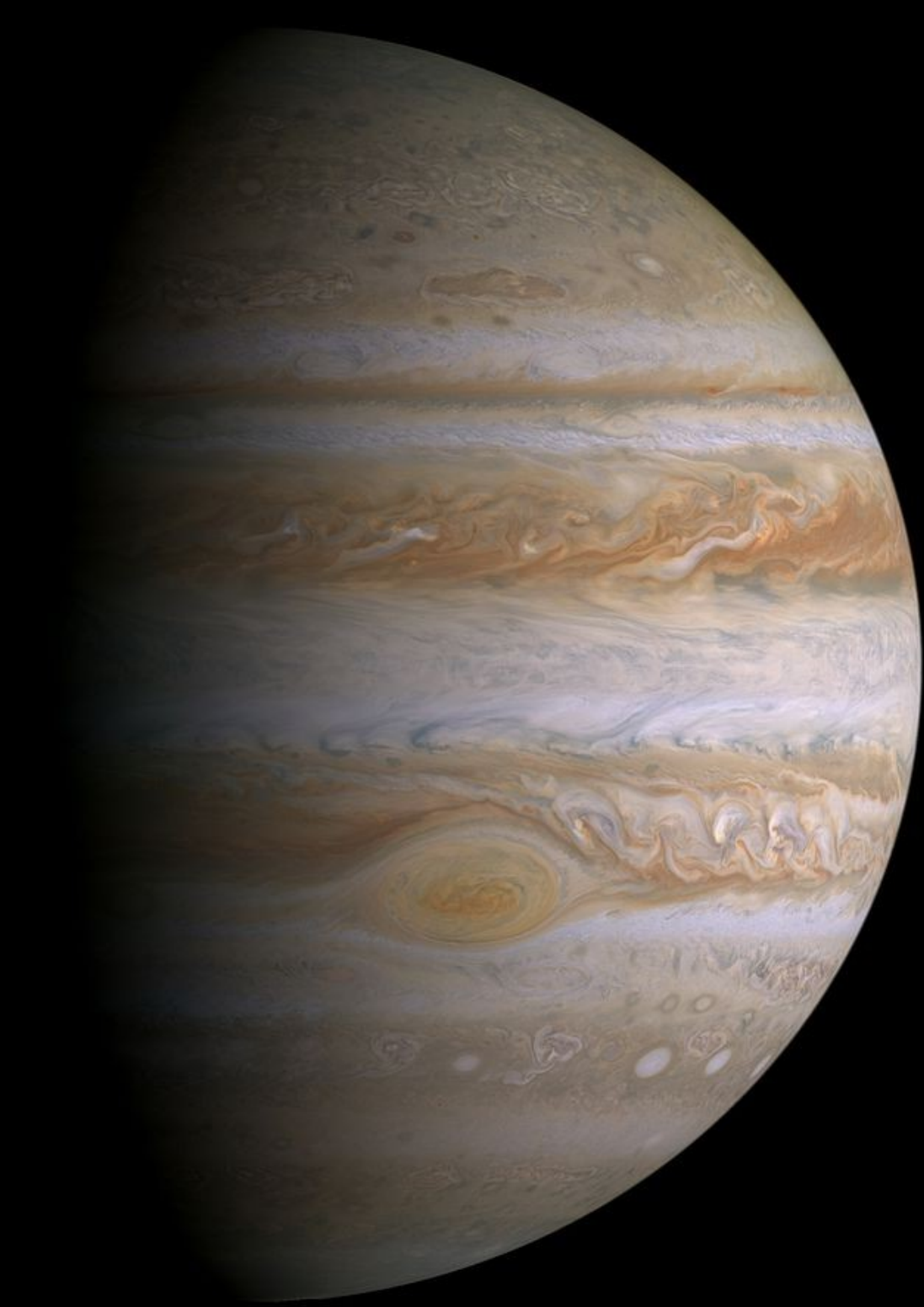
Tiny mass ( 1 )

Very tiny mass  
( $<0.00000001$ )









**x 150**

Electron  
Neutrino,  $\nu_e$



Electron  
Antineutrino,  $\bar{\nu}_e$

Muon  
Neutrino,  $\nu_\mu$



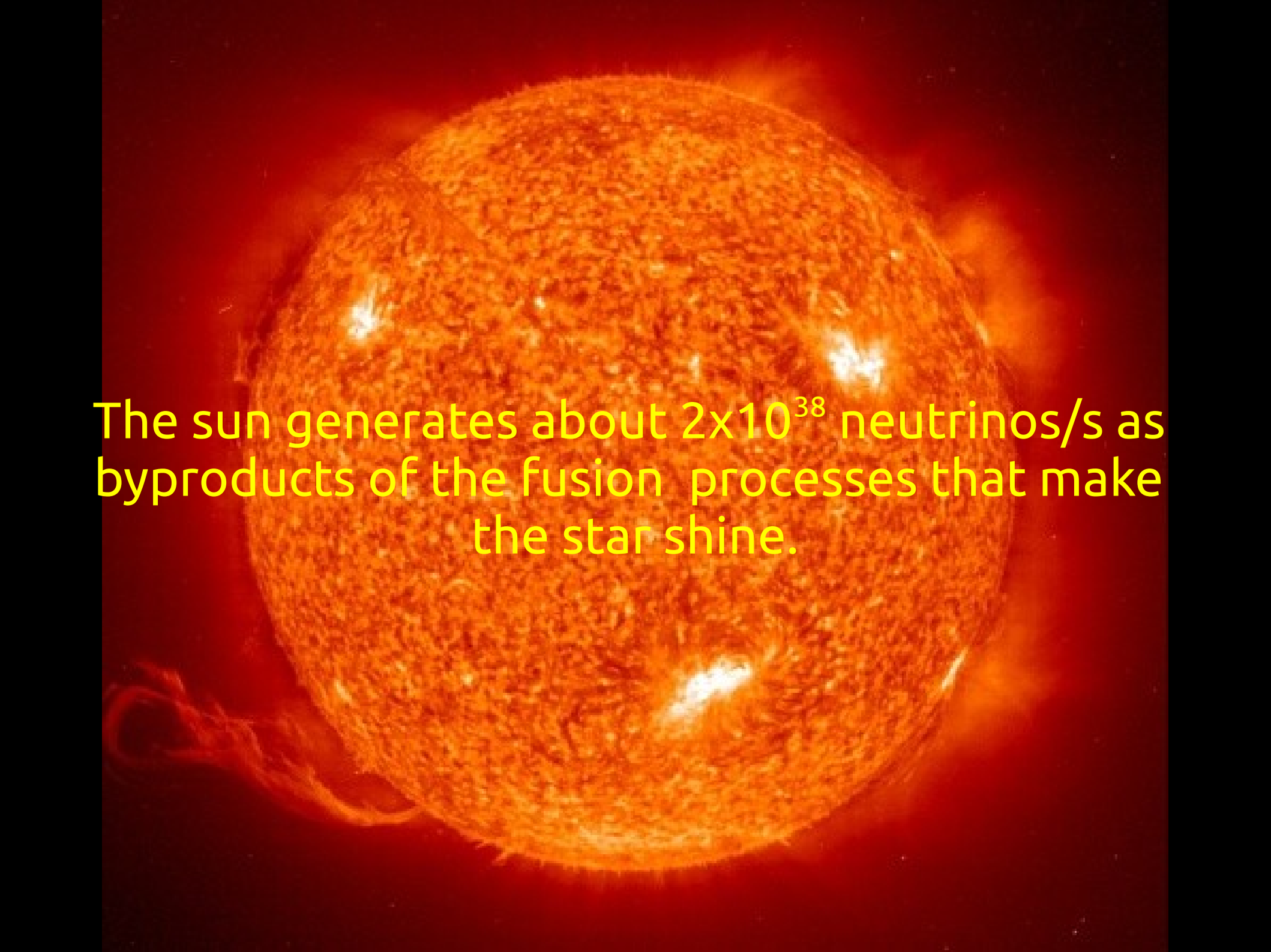
Muon  
Antineutrino,  $\bar{\nu}_\mu$

Tau  
Neutrino,  $\nu_\tau$



Tau  
Antineutrino,  $\bar{\nu}_\tau$

*3 neutrino Flavours*




The sun generates about  $2 \times 10^{38}$  neutrinos/s as byproducts of the fusion processes that make the star shine.



# So why don't we notice?

$\nu$  are almost ghosts. They interact extremely weakly with matter.

To a neutrino a planet is mostly empty space.



500,000,000,000,000 neutrinos  
from the sun just went through  
each and every one of you

"The chances of a neutrino actually hitting something as it travels through all this howling emptiness are roughly comparable to that of dropping a ball bearing at random from a cruising 747 and hitting, say, an egg sandwich."

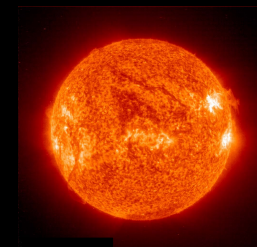
Douglas Adams - Mostly Harmless

Probability  $\approx 1 \times 10^{-13}$



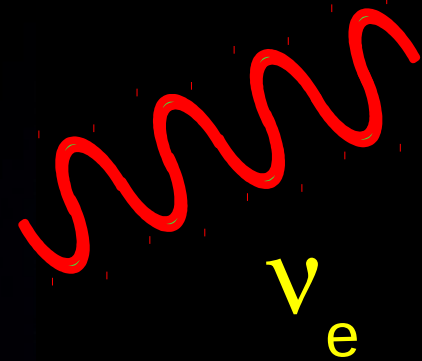
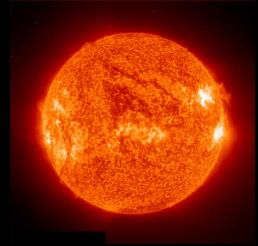
Egg Sandwich





$\nu_e$

Probability  $\approx 5 \times 10^{-13}$



How do we use neutrinos to study the matter/anti-matter asymmetry?

Electron  
Neutrino,  $\nu_e$



Electron  
Antineutrino,  $\bar{\nu}_e$

Muon  
Neutrino,  $\nu_\mu$



Muon  
Antineutrino,  $\bar{\nu}_\mu$

Tau  
Neutrino,  $\nu_\tau$

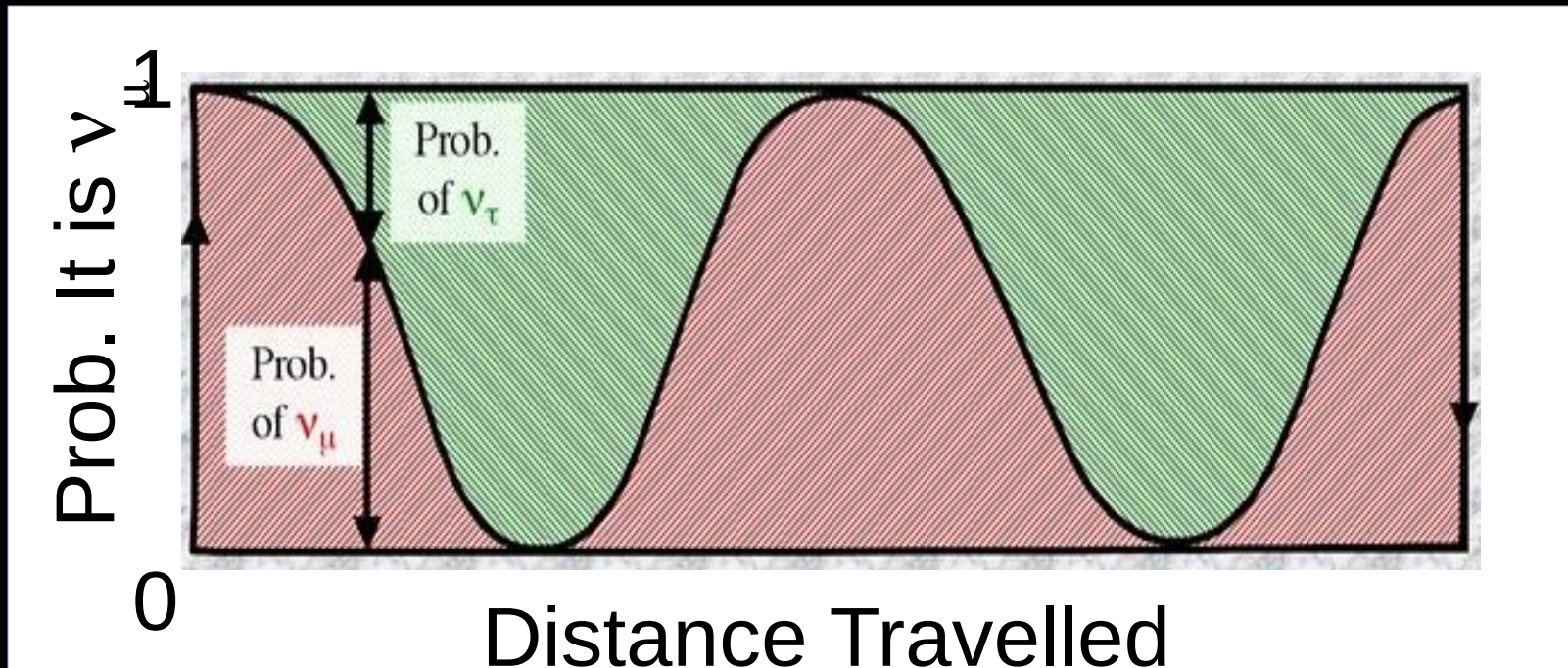
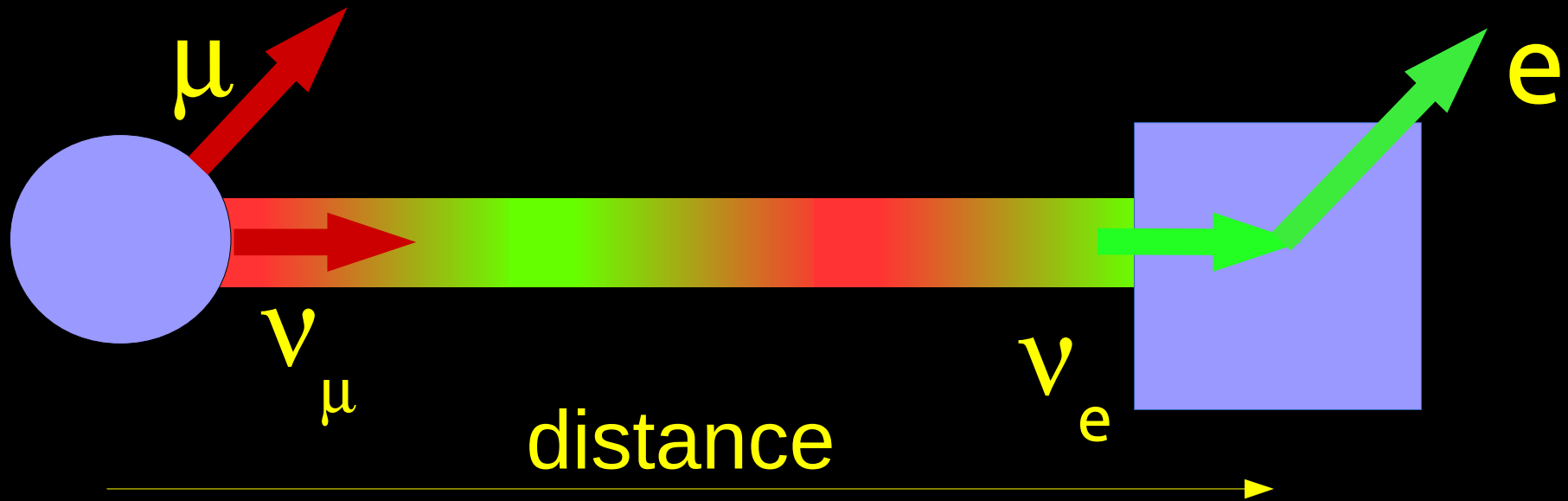


Tau  
Antineutrino,  $\bar{\nu}_\tau$

*3 neutrino Flavours*



# Neutrino Flavour Oscillations



## How do we use this?

$$Prob(\nu_{\mu} \rightarrow \nu_e) \neq Prob(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)$$

Then neutrinos behave differently from anti-neutrinos

An idea floating around suggests that, if this happens, then the same thing will happen to matter and anti-matter

# T2K Experiment



90%  $\nu_\mu$



Super-Kamiokande

$\mu$



295 km



JPARC

10%  $\nu_e$

100%  $\nu_\mu$

$\mu$

Image © 2008 TerraMetrics

Image NASA

Image © 2008 Digital Earth Technology

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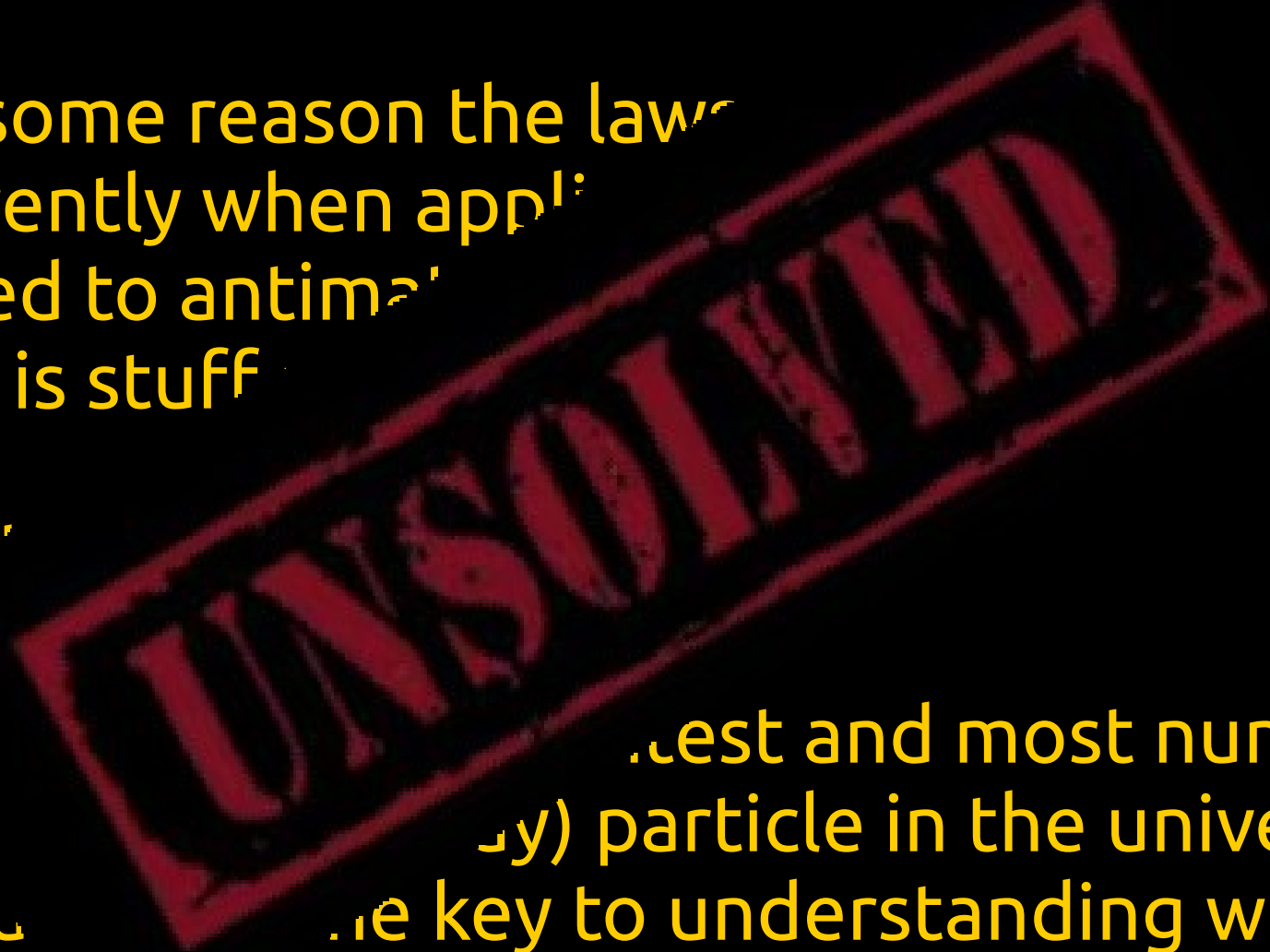


# The case of the missing antimatter

- For some reason the laws of physics behave differently when applied to matter than when applied to antimatter. This is the reason why there is stuff around at all.
- We don't know why (yet)
- The neutrino – the lightest and most numerous (but hardest to study) particle in the universe may just hold the key to understanding why we are here at all.

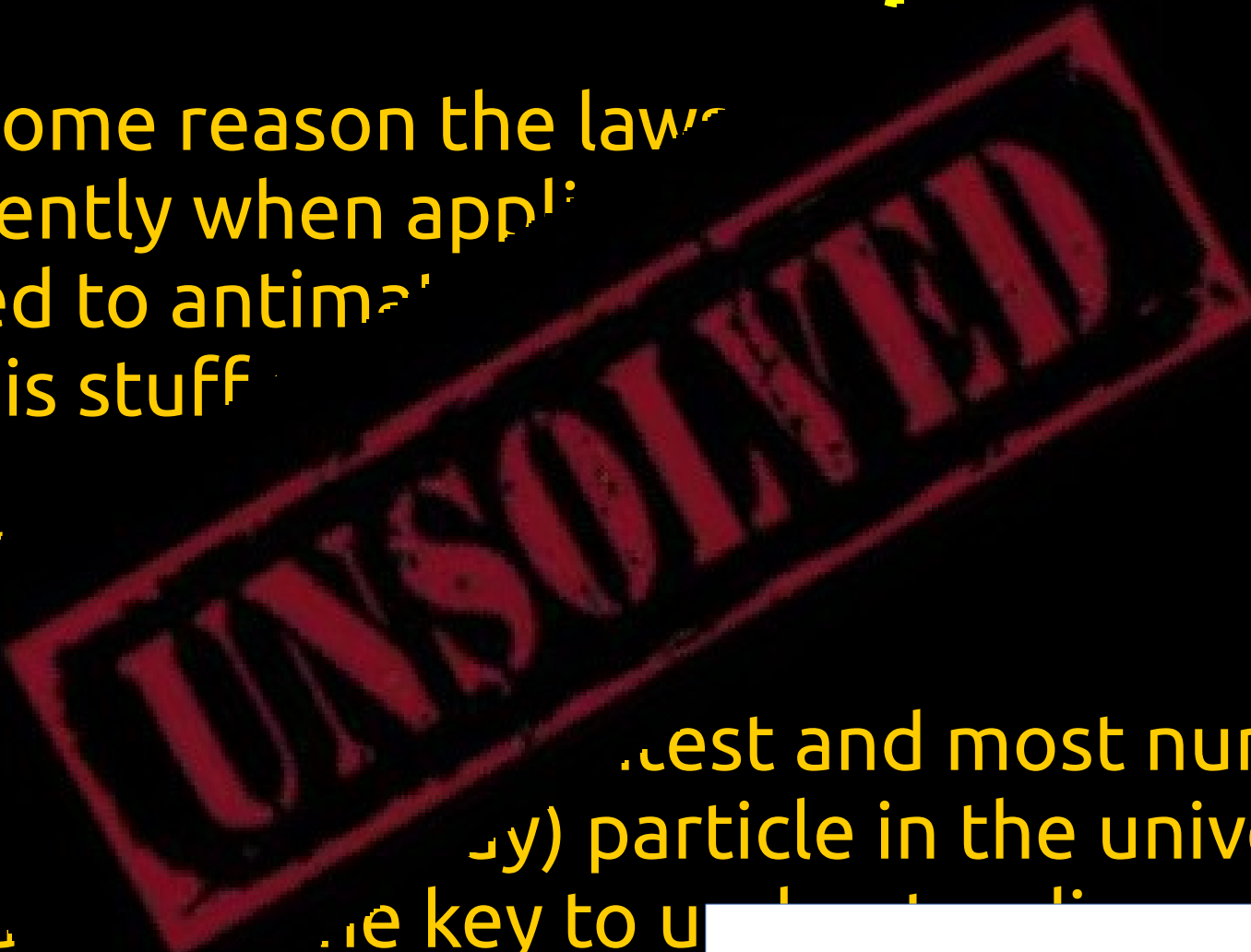
# The case of the missing matter

- For some reason the laws of physics seem to behave differently when applied to antimatter than when applied to matter. There is stuff missing.
- We know that dark matter is the most abundant and most numerous (but we don't know what it is) particle in the universe and it may just be the key to understanding why we are here at all.



# The case of the missing matter

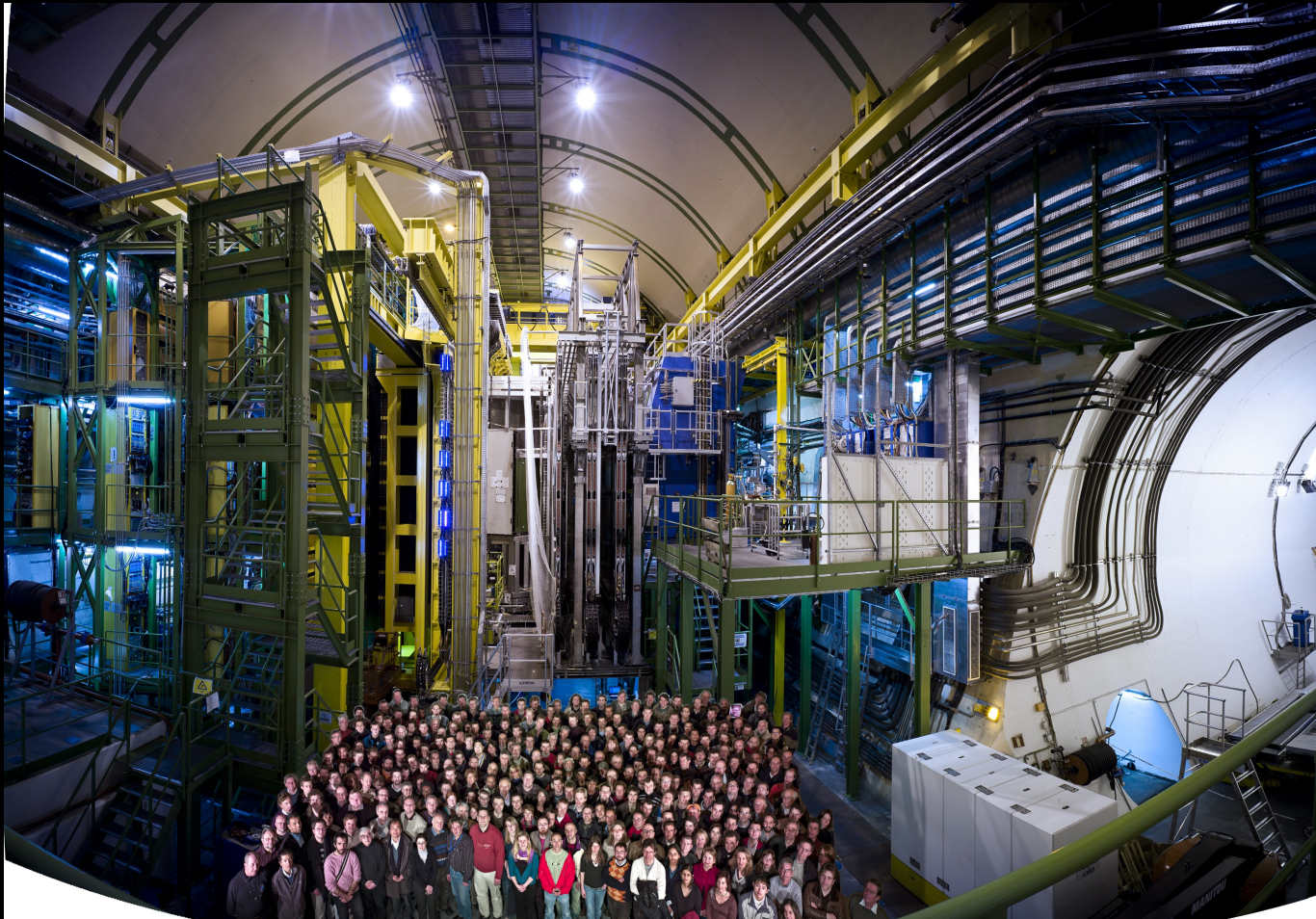
- For some reason the laws of physics seem to behave differently when applied to antimatter than when applied to matter. There is stuff missing.
- We know that dark matter is out there.
- The most abundant, lightest and most numerous (but very hard to detect) particle in the universe may just be the key to understanding the missing matter. It's all here.



So far.....

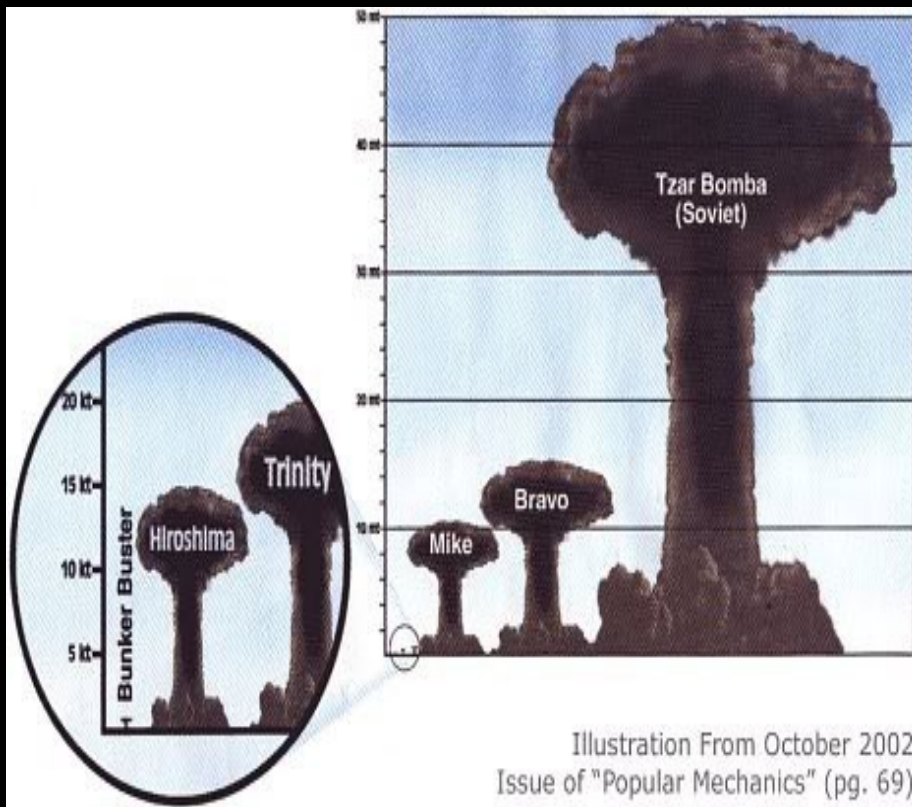
There *is* a difference between the physics of matter and antimatter. It's name is *CP Violation*

LHCb



The LHC will study this by looking differences between particles called  $B^0$  and  $\bar{B}^0$  mesons

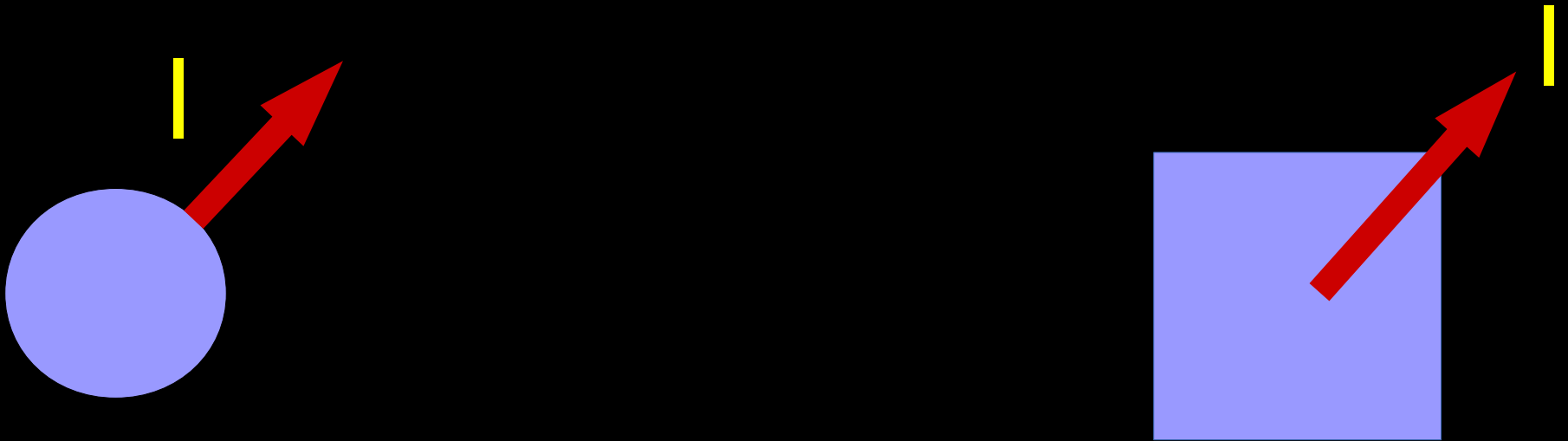
1 kilogram of matter  
+  
1 kilogram of antimatter





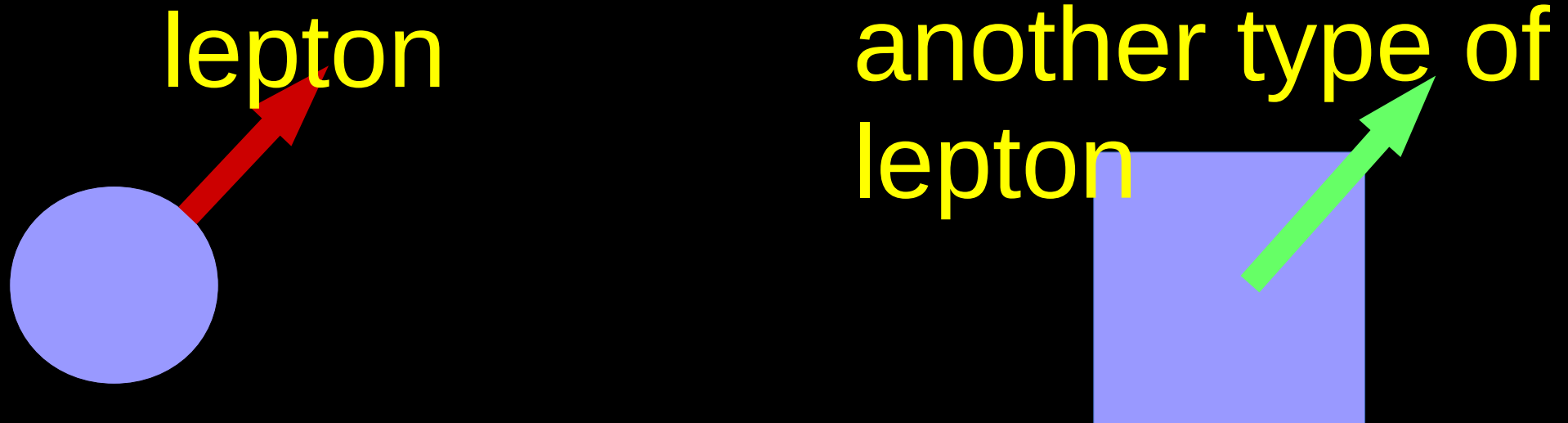
# Neutrino Oscillations

*THE* most important discovery about neutrinos in the last 20 years



A typical neutrino experiment

# Neutrino Oscillations



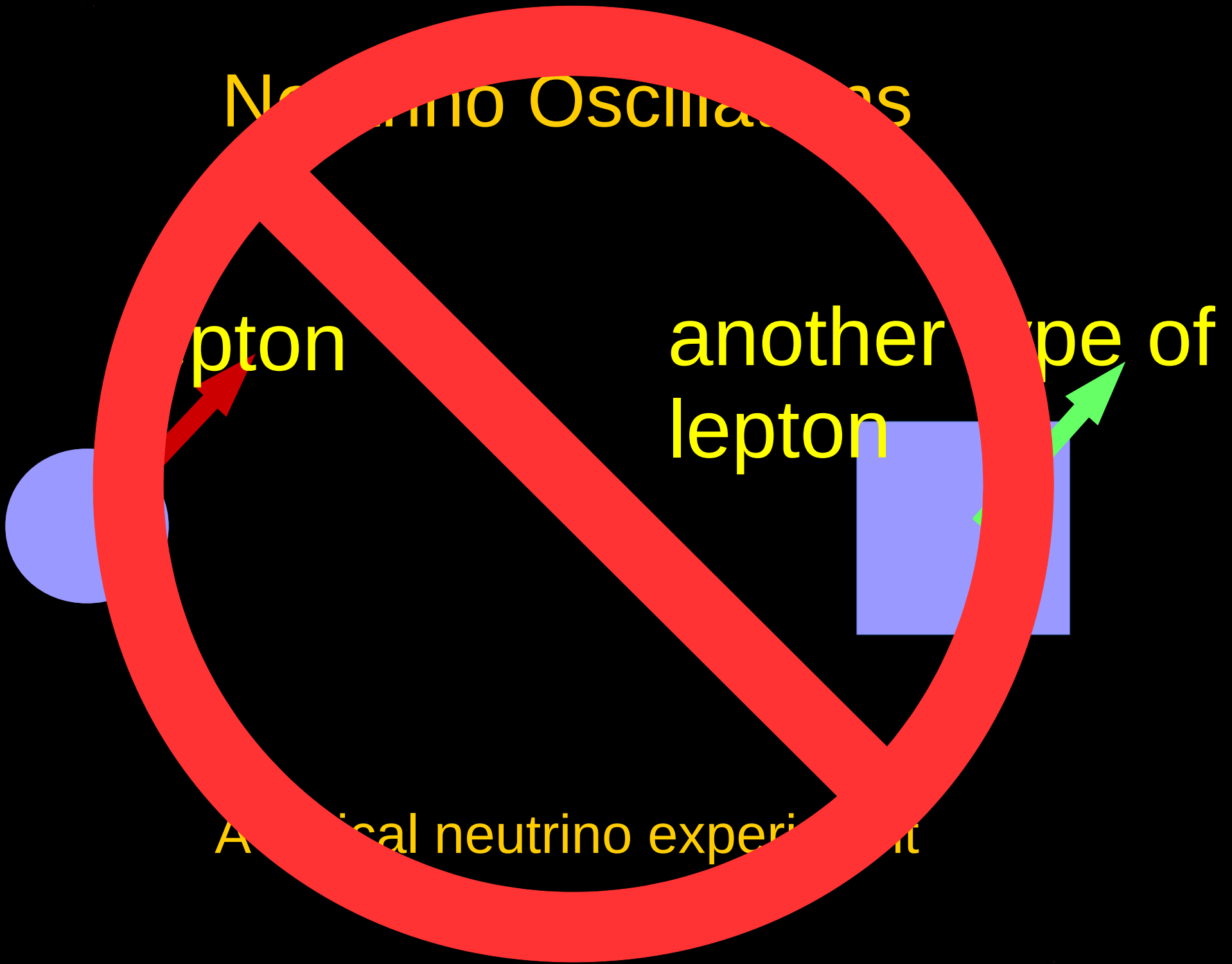
A typical neutrino experiment

# Neutrino Oscillations

muon  
neutrino

another type of  
lepton

A real neutrino experiment



# It's Quantum.....

$$|\nu_\alpha\rangle = \sum_{i=1}^3 U_{\alpha i} |\nu_i\rangle \quad \text{where } U_{\alpha i} \text{ is a unitary matrix}$$

$$|\nu_k(t, x)\rangle = e^{i(E_k t - p_k x)} |\nu_k(0,0)\rangle \rightarrow P(\nu_\alpha(0,0) \rightarrow \nu_\beta(t, x)) = |\langle \nu_\beta(t, x) | \nu_\alpha(0,0) \rangle|^2$$

$$|\langle \nu_\beta(t, x) | \nu_\alpha(0,0) \rangle|^2 = \sum_k \sum_j U_{\alpha k} U_{\alpha j}^* U_{\beta k} U_{\beta j}^* e^{i((E_j - E_k)t - (p_j - p_k)x)}$$

$$(E_j - E_i)t - (p_j - p_i)x = (\sqrt{p_j^2 + m_j^2} - \sqrt{p_i^2 + m_i^2})x - (p_j - p_i)x =$$

$$\left( p_j \left( 1 + \frac{1}{2} \frac{m_j^2}{p_j^2} \right) - p_i \left( 1 + \frac{1}{2} \frac{m_i^2}{p_i^2} \right) \right) \approx \frac{\Delta m_{ij}^2}{4E}$$



$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \longrightarrow P(\nu_\alpha(0,0) \rightarrow \nu_\beta(t, x)) = \sin^2(2\theta) \sin^2\left(\frac{\Delta m_{12}^2 L}{4E}\right)$$

# Positron Emission Tomography (PET)

